

Historic Great comets

1600 -2010

Great comets – Ancient observations / theories

Have inspired civilizations over millenia

Ancient Greeks

-Thought that comets were atmospheric objects
- (Aristotle 3rd cent BC)

-Ancient Romans

-Believed comets were celestial in nature(Seneca 1st AD)

-Europe – maintained Aristotelian view of cosmos into
- Medieval times

-Tycho Brahe/ Kepler – proved comets are in orbit around Sun



Wood-cut from German Pamphlet (Great comet of 1556)

What is a Historic Great Comet ?

3 main properties for great comet :

(a) **Closeness to the Earth**

(1861 Great comet : 0.82 au from Sun / 0.13 au from Earth)

- visible for many hours in dark sky / large developed dust tail

(b) **Closeness to the Sun**

(Kreutz sungrazers : over 0.8au from Earth / 0.05au from Sun)

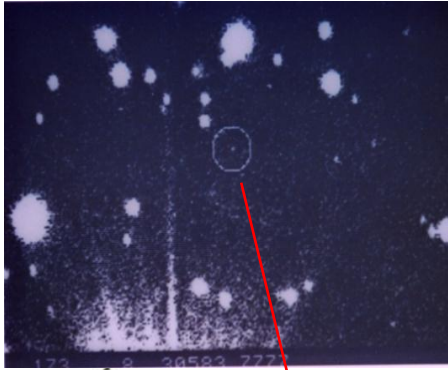
- extremely conspicuous in close perihelion passage

(c) **Very large and active**

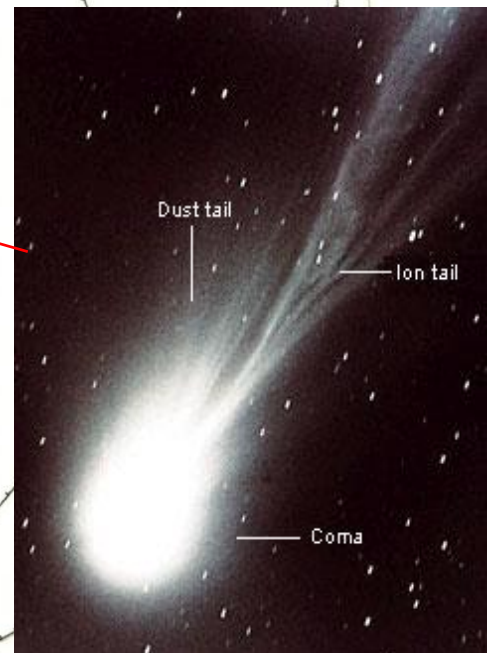
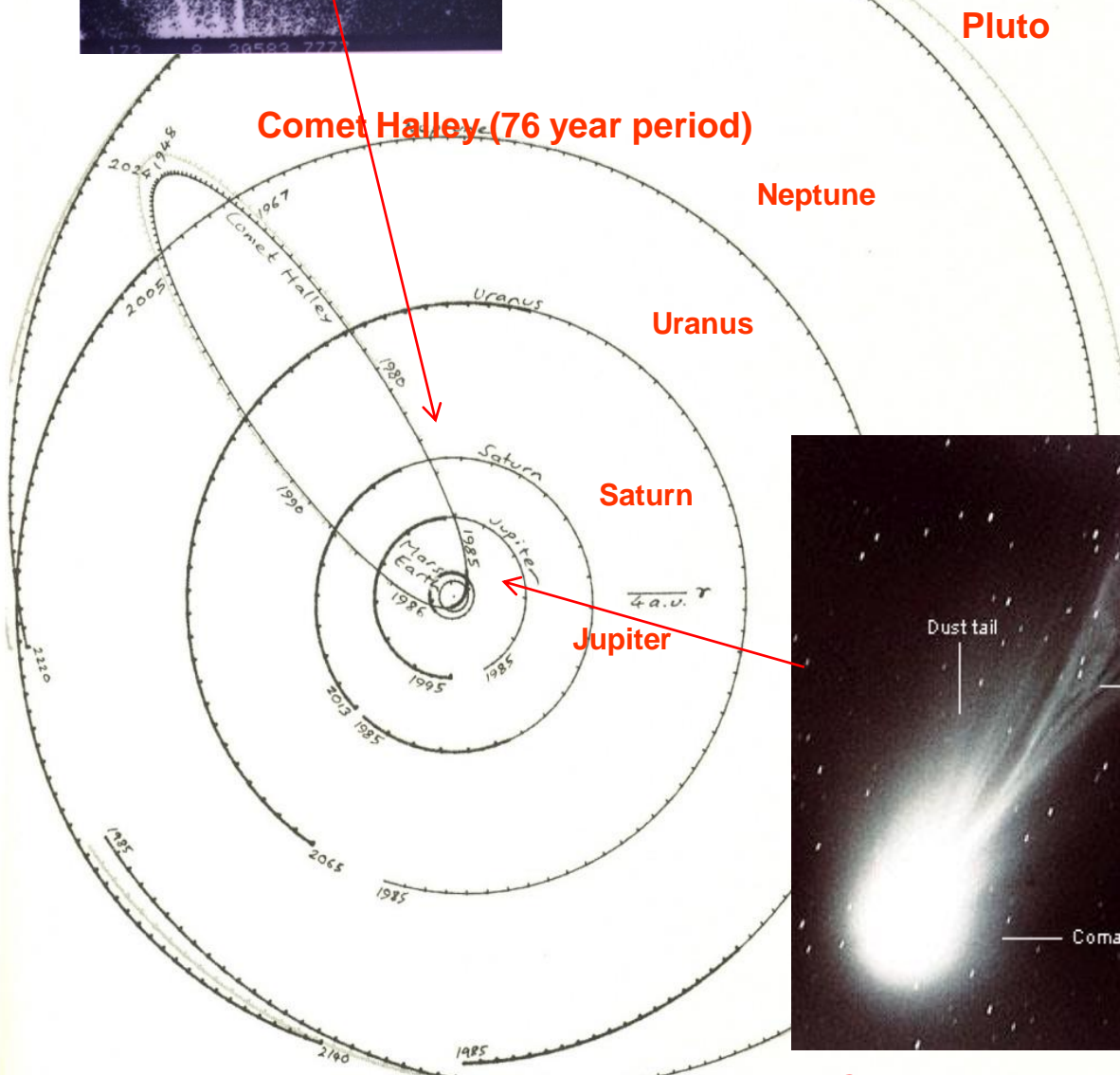
(Great comet of 1811 / Hale Bopp)

- quite rare for historic great comets
- naked eye visibility for over 100 days (260 days for 1811 comet)

Types of comet orbits – Short Period



Comet Halley recovery image (Oct 1982)
5.1m Palomar telescope / CCD camera
Magn +24.2 (bright star at bottom Magn+21)
Comet 11Au from Sun



The entire orbit of Halley's Comet among the planets.

Comet Period < 200 years

Comet Halley showing
Developing coma/gas
& ion tails

Types of Comet Orbit – Long Period

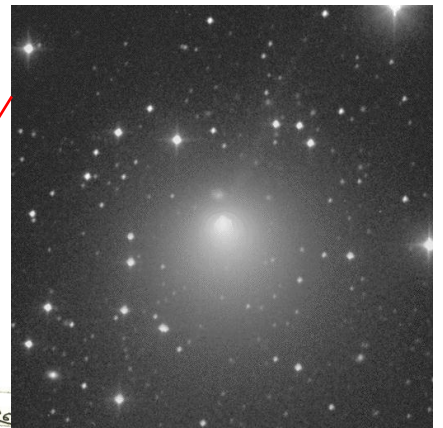
Period > 200 years –

Returning comets (eg Hale Bopp) – 4300 year period

Fresh (dynamically new) Oort comet – 100,000 year period



Comet Ikeya – Seki close to sun
(long period comet)



Comet Encke
(short period comet)

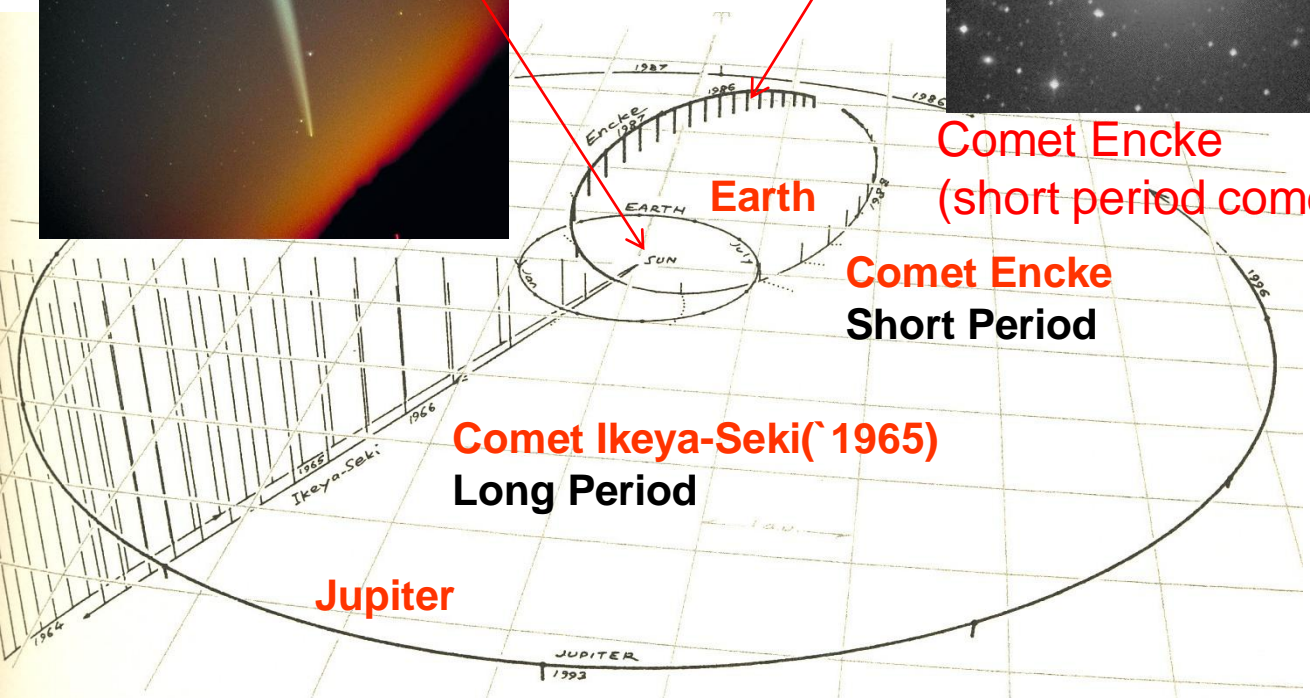


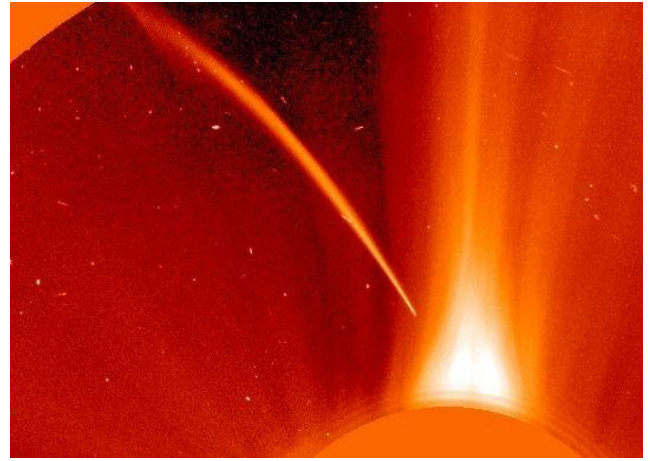
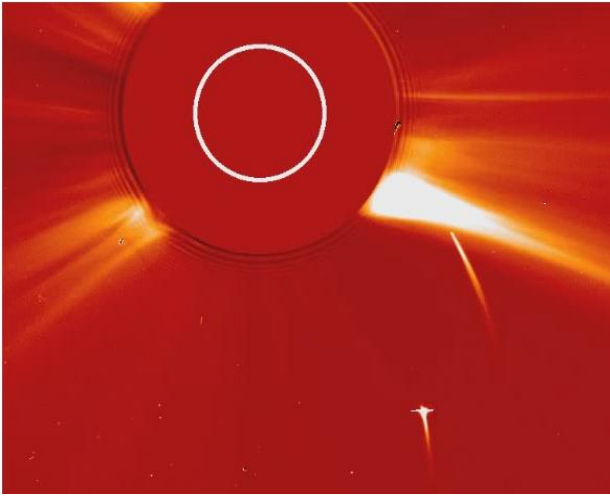
Diagram of the orbits of Comet Ikeya-Seki (long-period comet) and Comet Encke (short-period comet).

Sungrazer Comets

Close perihelion passage to Sun (within few thousand km)

Influence of Sun – Evaporation of comet ices/ tidal fragmentation

Examples of Sungrazing comets(SOHO) :



SOHO images

Kreutz sungrazing comets :

--Fragments of one giant comet – great comets of 1843/1882/1965 (75 mile diameter nucleus- fragmented 10,000 year ago)

--plunge into sun / disintegrate during perihelion passage

---usually are high inclination comets – perihelion < 2AU

Gravitational perturbations – reduce perihelion to low level

--- future great comet – likely to be Kreutz Sungrazer

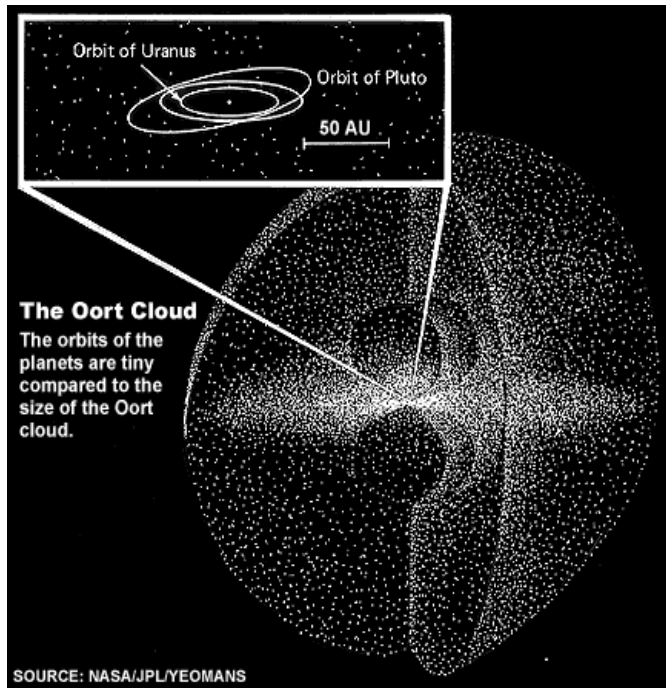
----- Great comet of 1680 –sungrazer – Newton verified Kepler`s

equations of orbital motion

The Origin of Comets - Oort Cloud

Immense spherical cloud surrounding solar system

Comets within Oort cloud –weakly held by Sun- perturbed by nearby stars
(solar influence out to about 2 light years)



Oort Cloud –40xMass of Earth

Source of periodic comets

Average distance –
44,000AU from sun

Discovered by Jan
Oort(1950)

Comets in Oort cloud :

- Icy balls of H_2O / CH_4 / NH_3
- & rock /dust

Oort Cloud formation :

Formed 4.5 billion years ago from icy/rocky debris (planetesimals)

Planetesimals ejected into large orbit by gravitational interaction with Jupiter

Ejected debris either escaped from solar system / populated Oort cloud

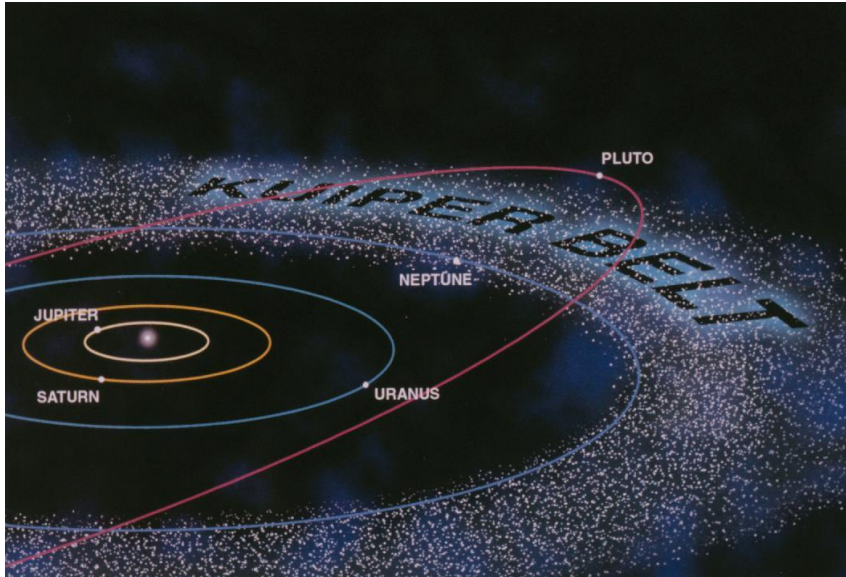
Comets from Oort Cloud:

Gravitational interaction between comet & passing star or molecular cloud

Causes comet to fall into solar system on long period elliptical orbit

The Origin of Comets – Kuiper Belt

Reservoir of dormant icy comets just beyond Neptune (30- 1000AU dist)



Orbital features:

**Orbital direction same
As planets**

Slightly inclined orbits

**Source of short period
comets**

1110 Kuiper belt objects known (10-50km) – millions more smaller than 100m

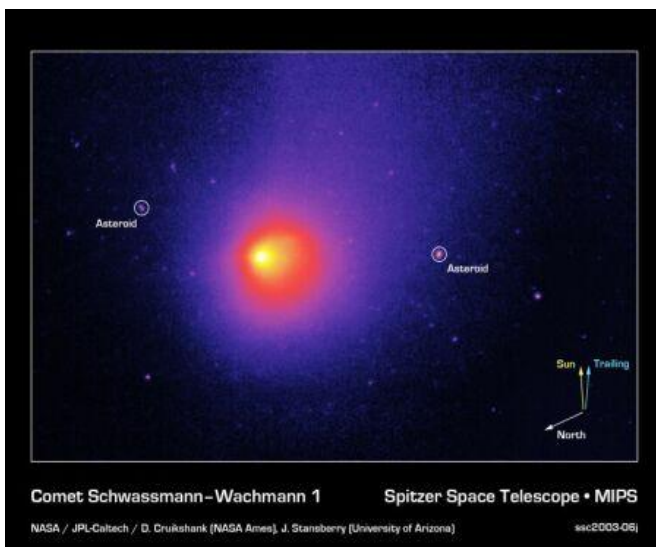
Kuiper belt object characteristics

icy pristine objects – mainly water / methane

Gravitational interaction with Neptune / Uranus –

- (i) eject Kuiper belt object into Oort Cloud / Interstellar space

(ii) send object into inner solar system as **short period comet**
(unstable orbit)



Kuiper Belt objects:

Show Presence of coma

-Chiron (170km diameter)

-20x size of Halley comet nucleus

-Potential spectacular comet

Comet Structure – the Nucleus

Irregular shaped body – about 10 to 20 km in size

Nucleus composition – ices{H₂O, CO₂, NH₃} / stony & metallic solids surrounded by dark crust - **dirty snowball**

Dormant in outer solar system

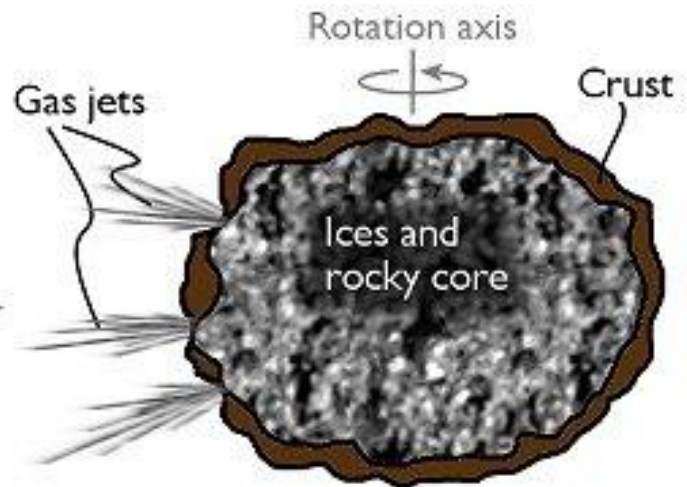
Close to Perihelion :

(1) Gases evaporate in jets

Venting through crust 

(2) Material in jets form coma

and comet tails



Surface of Comet nucleus :

Dark material (albedo of 4%)

Dust / complex organic molecules

Temperature of outer layers of 330K

Gas jets generate

Non-gravitational forces
(influence on comet orbital period)

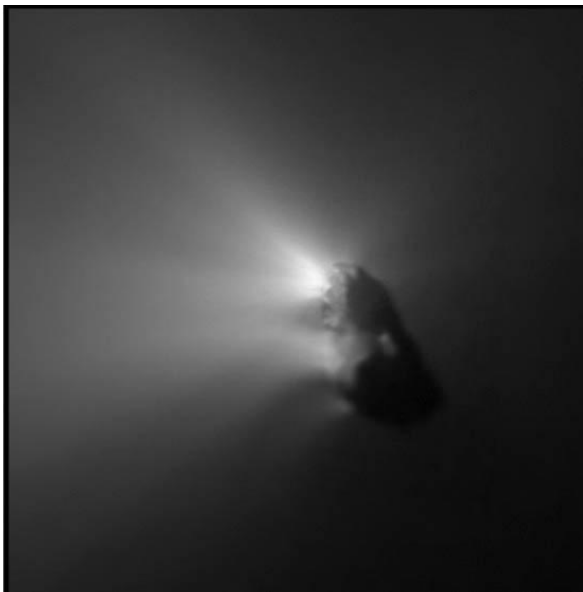
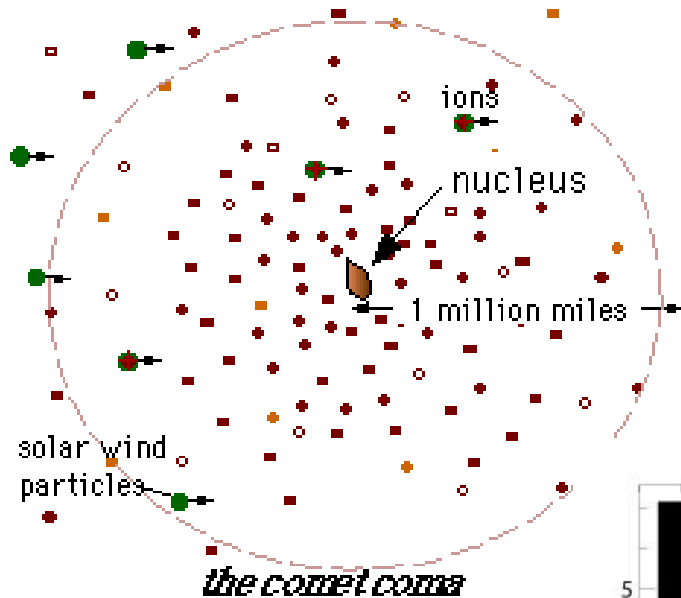


Image of Halley`s Comet
(Giotto Probe 1986)

Complex rotation of nucleus – rotation / spin axes and precession

Comet Structure – the Coma

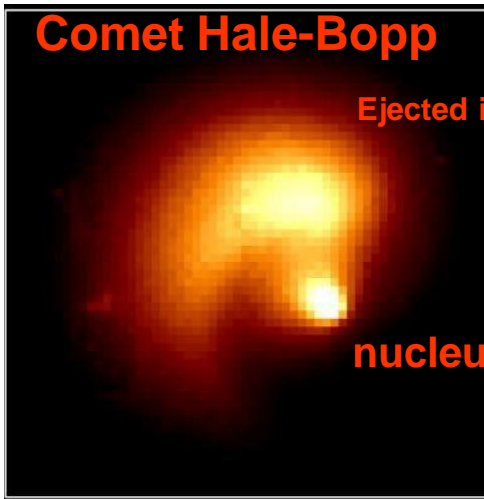
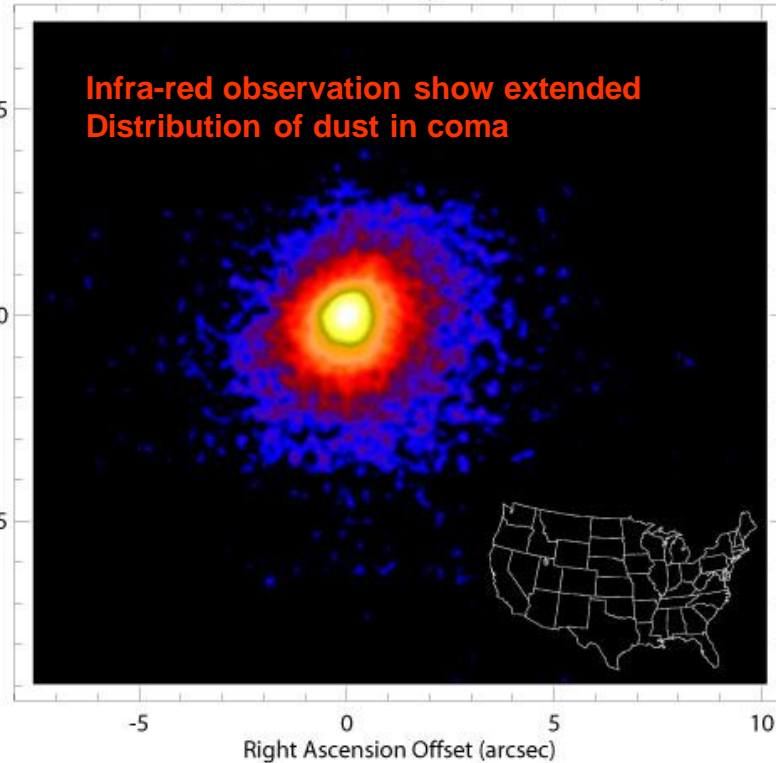
Cometary ice evaporation – formation of atmosphere around nucleus (coma)



Coma characteristics :

- Very low density
- Millions of miles in extent
- Neutral Coma particles excited by solar wind and ionized .

Comet Tempel 1 - 2005 July 6 01:00 UT - 11.7 μm



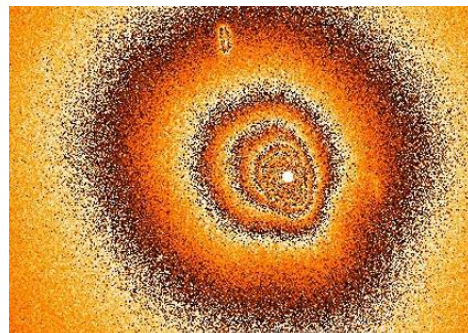
Hubble image showing :

Ejection of icy crust from nucleus due to evaporation/ comet rotation

Distribution of ice particles around Nucleus in coma

Cloud of material in coma is 3.5x Fainter than nucleus

Image of Central part of Hale Bopp nucleus :



Distribution of molecules / dust

Fan / Jet structures in coma

ESO observatory

Comet Structure – Gas / Ion Tails

Components Of Comets

Dust / ion tails Comet tails form as Comet approaches the Sun

There are 3 features in the comet structure :

Hydrogen Envelope

-Hydrogen molecules absorb ultra-violet Light and form large envelope (invisible from Earth)

Ion Tail (Type I)

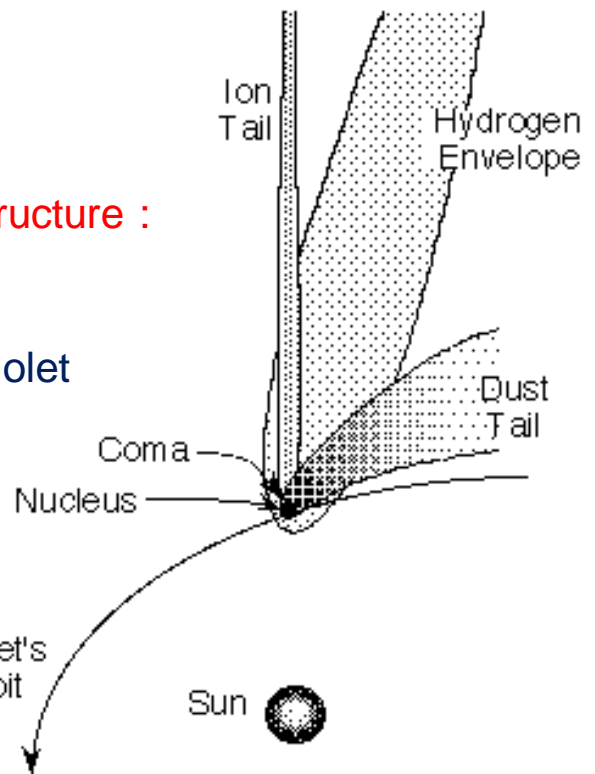
Low mass ions released from coma are accelerated away from Nucleus – directly away from Sun
Strongly affected by solar wind-

Dust Tail (Type II)

The dust particles are left behind along the comet's orbit and form a curved tail

-(carbon- rich /
-Silicate material) –

-(Mabe over 10 million
-Miles in length)

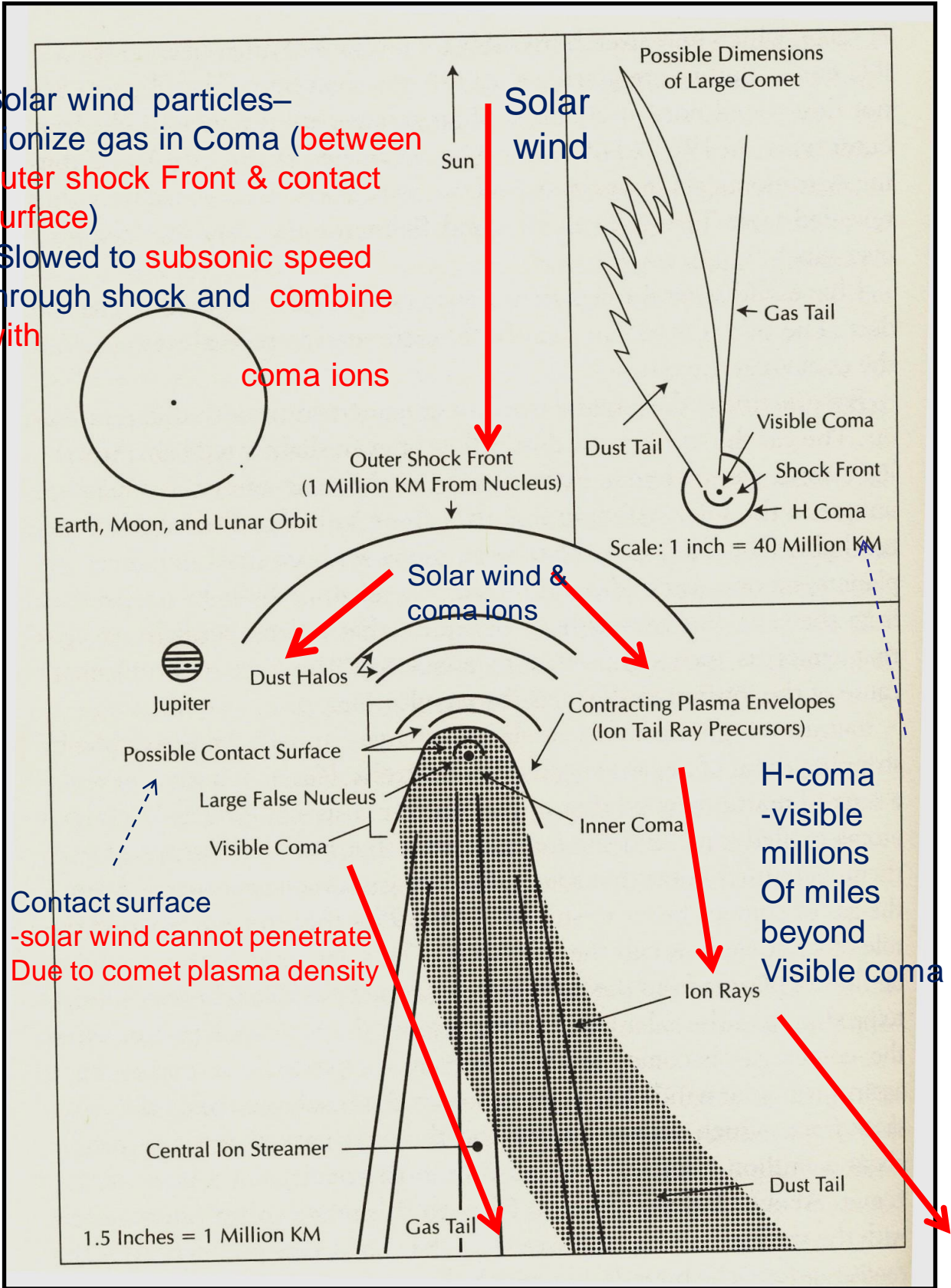


Comet P1
McNaught



Comet Structure : H Coma / Shock Fronts

Solar wind particles—
 - Ionize gas in Coma (between
 outer shock Front & contact
 surface)
 - Slowed to subsonic speed
 through shock and combine
 with
 - coma ions



Contact surface
 -solar wind cannot penetrate
 Due to comet plasma density

H-coma
 -visible
 millions
 Of miles
 beyond
 Visible coma

1.5 Inches = 1 Million KM

Scale: 1 inch = 40 Million KM

Halley's Comet through the Ages – 1404BC – 1682AD

Perihelion distance – 0.6AU)



Halley's comet (Bayeux Tapestry) 1066AD`

One of the most impressive returns of Halley` comet

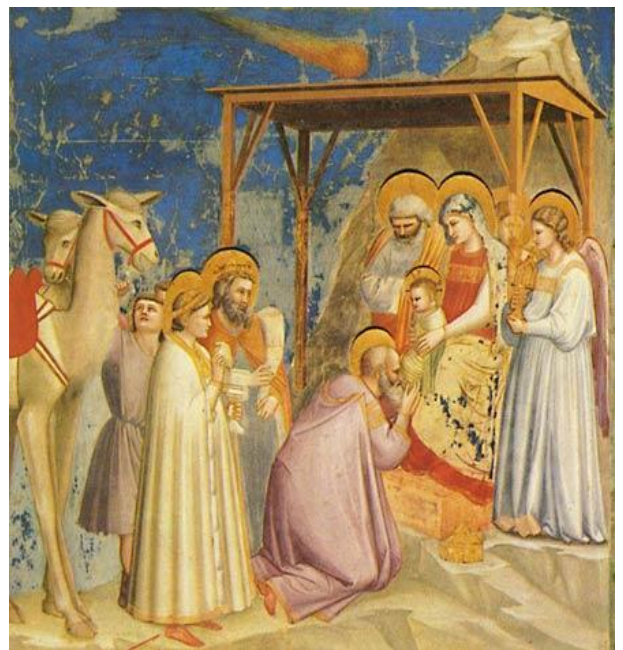
- Passed only 0.1au from Earth after Perihelion (magn-4)
- Tail spanning over 90deg (visible all night)

Halley`s comet (Adoration of Magi) → Giotto fresco

1301 (Giotto`s Christmas star)

pre-perihelion encounter
(comet only 0.18au from Earth)
Comet Magn+2 (tail 50-70deg)

-Giotto painted Nativity scene
-In Arena Chapel(Padua) in 1306
(realistic comet)



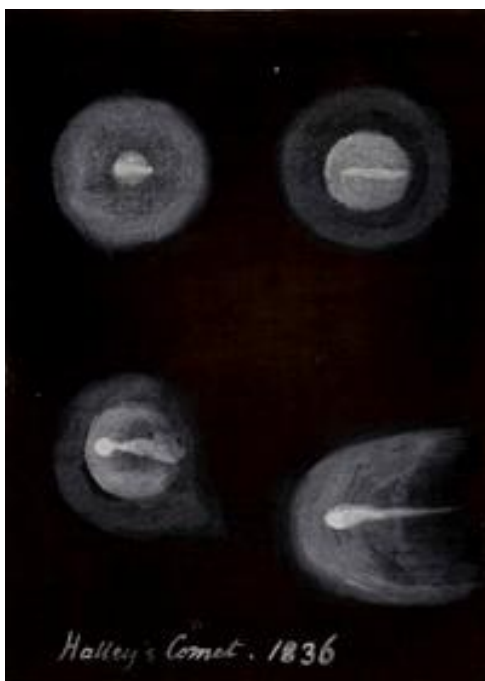
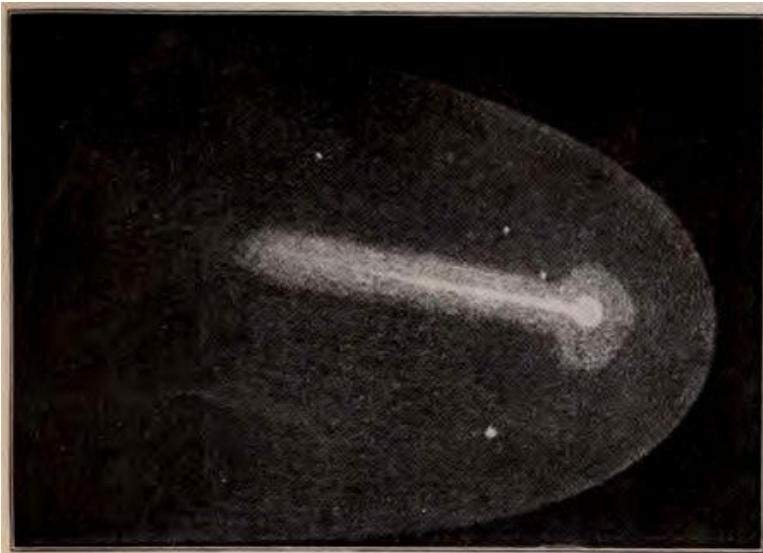
Halley's Comet through the Ages

1680AD – 1836AD

1682 Return – seen by Halley and Newton (Magn 0-1) 30deg tail

1759 – first predicted return – only 0.122au from Earth
(seen by Messier – best seen in southern hemisphere)

1835 – seen and sketched by Sir John Herschel (southern hemisphere)
fairly close (0.19au) to Earth before perihelion (Nov 1835)



After Perihelion – outburst of comet
1.5AU from Sun

-Brightened by 2 to 3 magnitudes

-- enormous halo(million miles in 3 weeks)

-Likened to transparent gauze

Halley's Comet through the Ages

1910AD



Close approach (0.15au)
& bright (magn -1.5)

Earth passed through gas tail
& leading edge of dust tail
missed by 240,000 miles

Last seen by Barnard on
May 23 1911 at 5 au from Sun

Spectroscopic analysis of tail
- Discovery of cyanogen

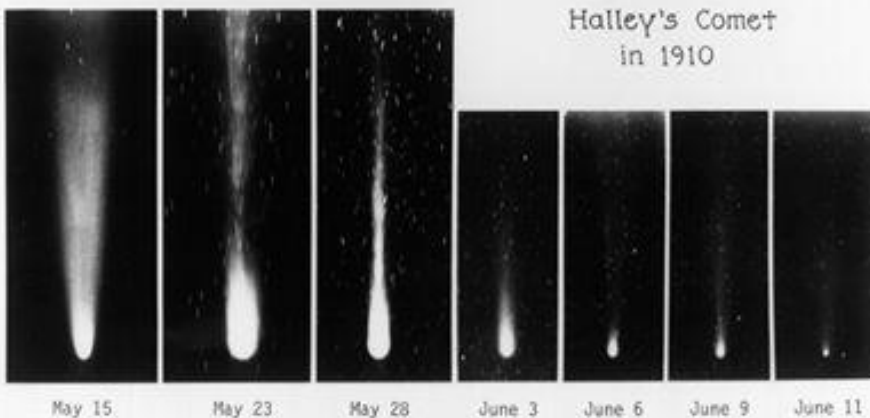
Disconnection events seen in tail –effect of Solar wind



Digitized / colourized image from 1910
(Kitt Peak Observatory)



Sequence of
Images from 1910
Showing
Comet Halley
(Apr1910-Jun1910)



Halley's Comet through the Ages

1986AD



Halley's comet(1986)

Halley's comet nucleus
(imaged by Giotto)



5 spacecraft sent to investigate Halley's comet :

Giotto (Jul 85) / Vega 1&2(Mar 86- USSR) / Sagikake & Suisei(Japan)

Halley Comet Events

CCD Recovery image(Mag24)
(Oct 16 1982 – 11 AU from Earth)

First visual sighting (Mag19.6)
(Jan 23 1985 – 4.3AU from Earth)

First Naked Eye sighting
(Nov 8 1985) – 0.9AU from Earth

Perihelion – Feb 9th 1986

Giotto Flyby - Mar 14 1986

Closest Approach to Earth
(Apr 1986)- 0.4AU (dust tail 30-50deg)



Ion Tail Detachment event
(Apr 1986 (Cerro Toledo Observatory))

Final naked eye sighting
(May 29 1986 – 1.7AU from Earth)

Great Comets (16th – 19th Century)

(a) 1577

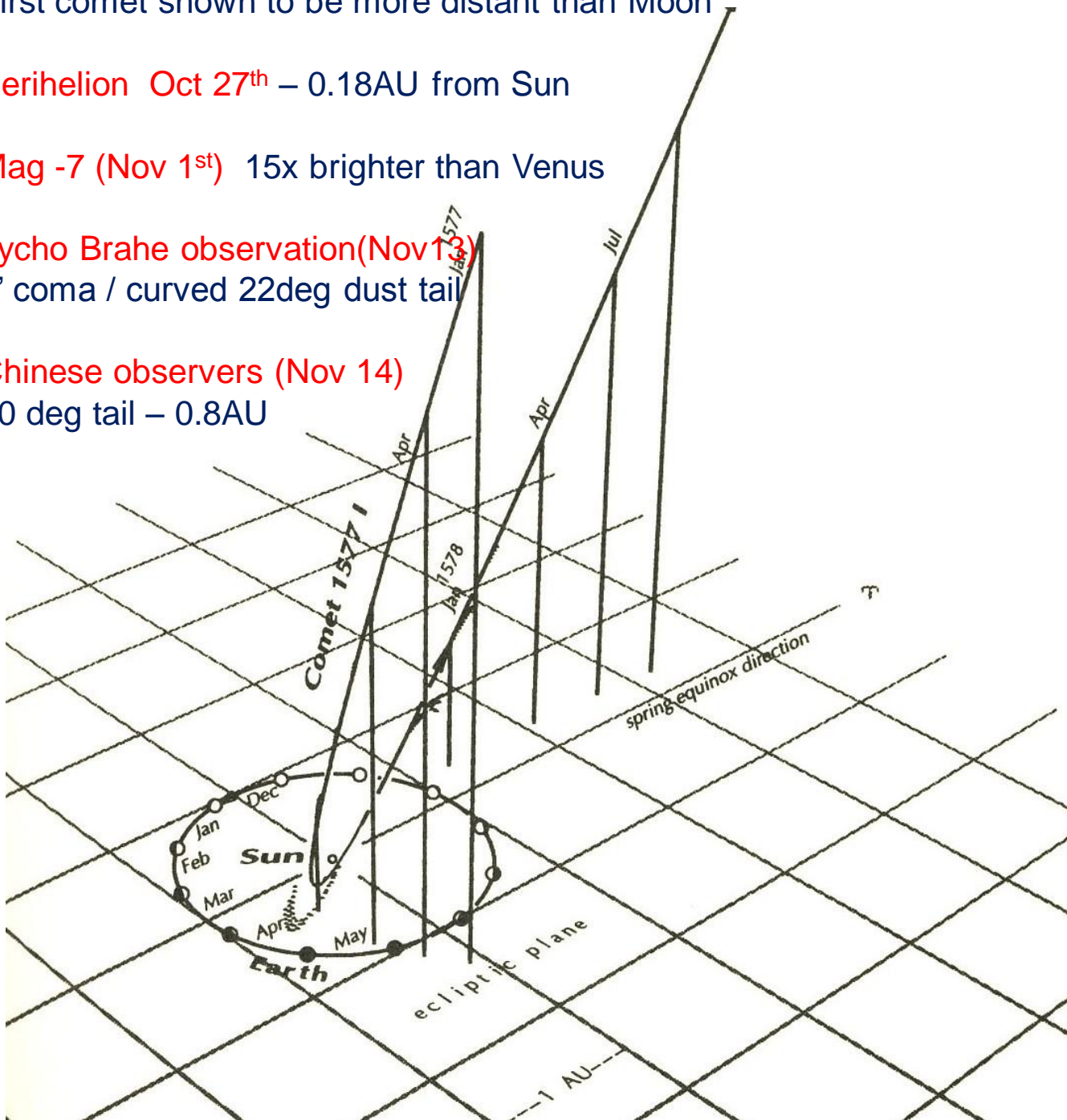
First comet shown to be more distant than Moon

Perihelion Oct 27th – 0.18AU from Sun

Mag -7 (Nov 1st) 15x brighter than Venus

Tycho Brahe observation (Nov 1st)
8' coma / curved 22deg dust tail

Chinese observers (Nov 14)
50 deg tail – 0.8AU



Most famous observer - Elizabeth I

Great Comets (16th – 19th Century)

(b) 1680

First great comet discovered with telescope

Perihelion : only (Earth-Moon) distance from Photosphere

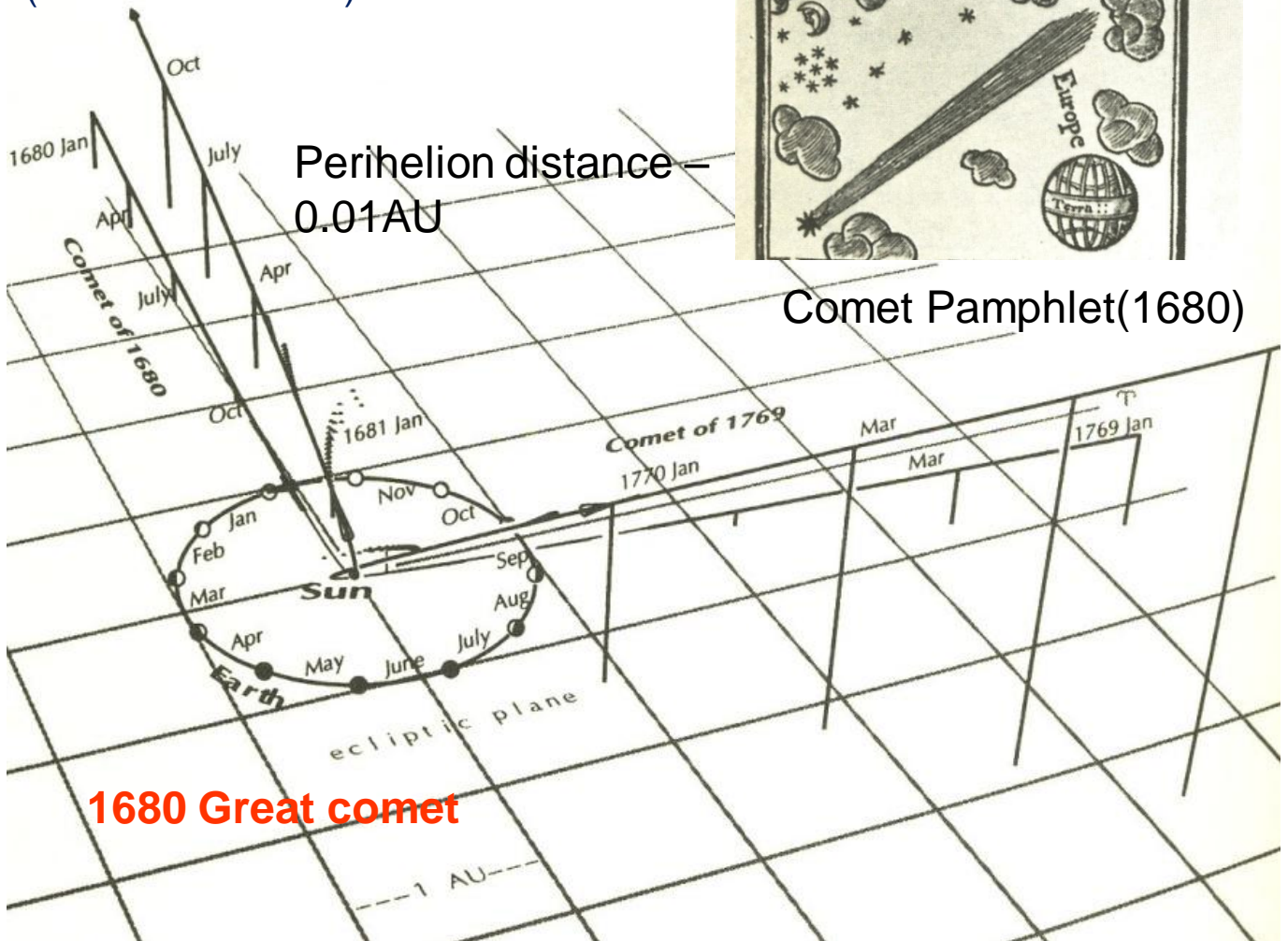
Famous observers:

Flamsteed (Dec20) – Greenwich
(tail halfway across sky – like sabre)

Issac Newton (Feb 18th 1681)
(2.2 au from Earth)



Comet Pamphlet(1680)



Great Comets (16th – 19th Century)

(c) 1744

Spotted at 1AU from Earth
(3 months before perihelion)

Feb 1 1744- brighter
than Venus

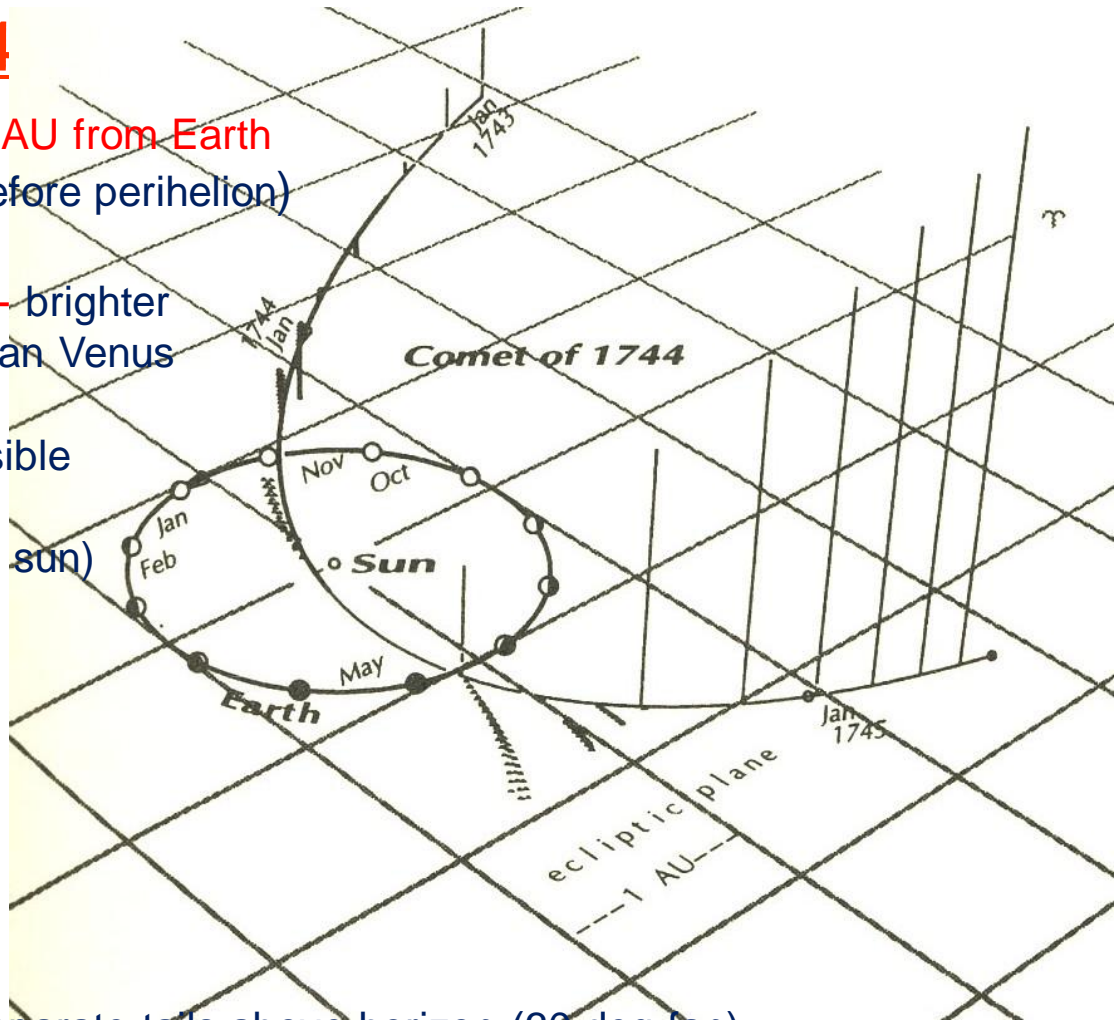
Feb 28 – visible
In daylight
(12deg from sun)

Mar 6 – 5 separate tails above horizon (30 deg fan)
(comet head 20 deg
below horizon)

Mar 18 – tail 90deg
Across sky

5 months of naked
Eye visibility

Greatest intrinsic
Brightness of any
sungrazer



Great Comets (16th – 19th Century)

(c) 1811

Found on Mar 25 1811

Huge / bright comet
(visible to naked eye longer than
Any comet in over 500 years)

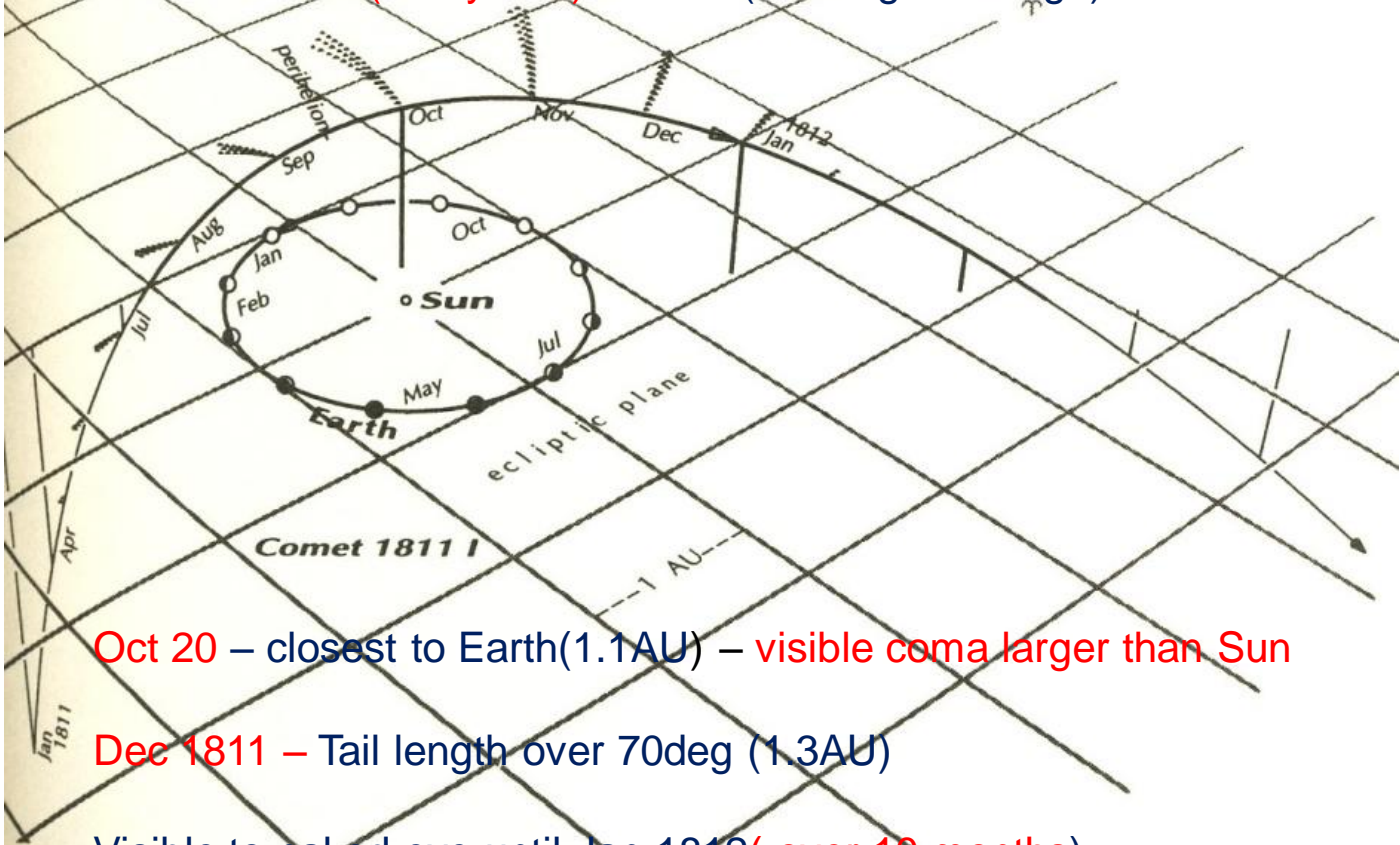
Perihelion distance of 1.04 AU

Oct 16 – circumpolar object
(gas tail 24deg)

Herschel – nucleus (ruddy hue) / coma (bluish/green tinge)



French caricature of the panic caused by the 1811 comet.



Oct 20 – closest to Earth(1.1AU) – visible coma larger than Sun

Dec 1811 – Tail length over 70deg (1.3AU)

Visible to naked eye until Jan 1812(over 10 months)

Seen in telescopes over 3.3AU from Earth

Great Comets (16th – 19th Century)

(d) 1843

The Great March comet
(Kreutz sungrazer)

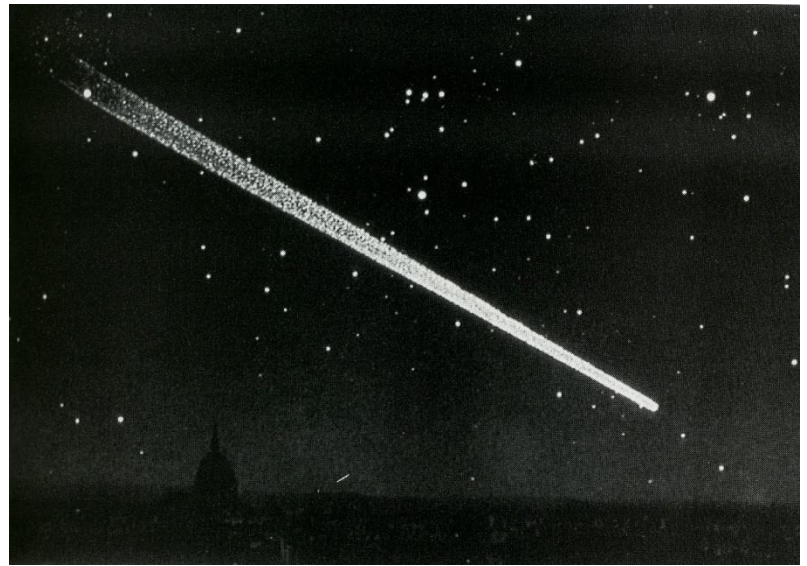
Spotted 1 deg from edge of sun
(broad daylight) –perihelion

Reported at Magn -7 and spotted
by Thousands of people

Tail of up to 90deg (Mar 4)
(190 million miles long)

Apparent nucleus of 12" and comaOf 45"

Mar 18th – comet seen far enough north to be seen in Europe (tail over 45deg)



Perihelion distance :
0.006AU

Great March Comet, 1843 I

Tail of over 64 deg (Mar 21)
Comet faded in April (over 2AU
From Earth

Orbit of Great Comet of 1843.

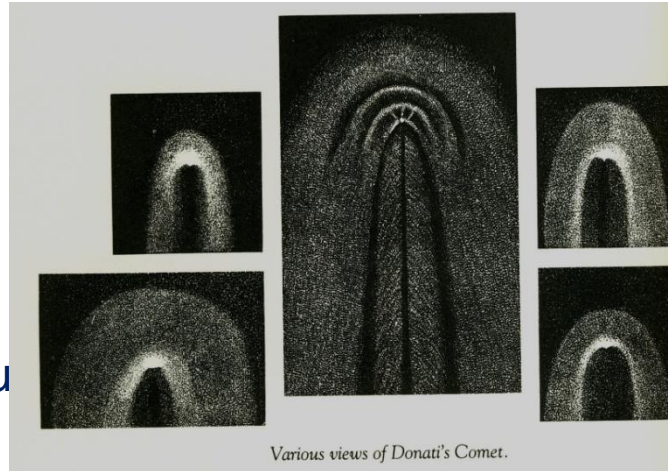
Great Comets (16th – 19th Century) 1858

Donati's Comet –

First visual sighting on Sept 6th

Perihelion reached on Sept 30th
(0.58au)

Interesting structure in nucleus
(false nucleus)



Comet head brighter than Arcturus
(magn -1) and tail over 35 deg – Oct 5

Broad dust tail –
Like a curved scimitar

2 delicate rays
- Twin components of gas tail



Donati's Comet on October 5, 1858.

Donati's Comet on October 9, 1858.

Comet closest to Earth –
Oct 10 (0.5au away/dust
Tail 60deg) - tail length 0.5au

Last telescopic view (Mar 1859)
- Magn 10

Orbit of the great comet of 1858 (Donati's).

Donati's Comet



Donati's comet as seen from Paris (1858)

Great Comets (16th – 19th Century)

(d) 1861

Closest great comet (0.13au)

Broadest / longest fan tail
(80deg fan & 50deg tails)

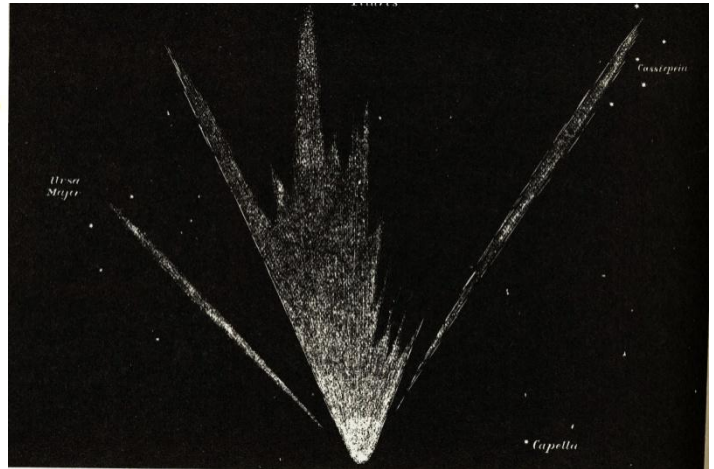
Highly inclined orbit (86 deg)

Earth passed through tail
(0.34 au in length)

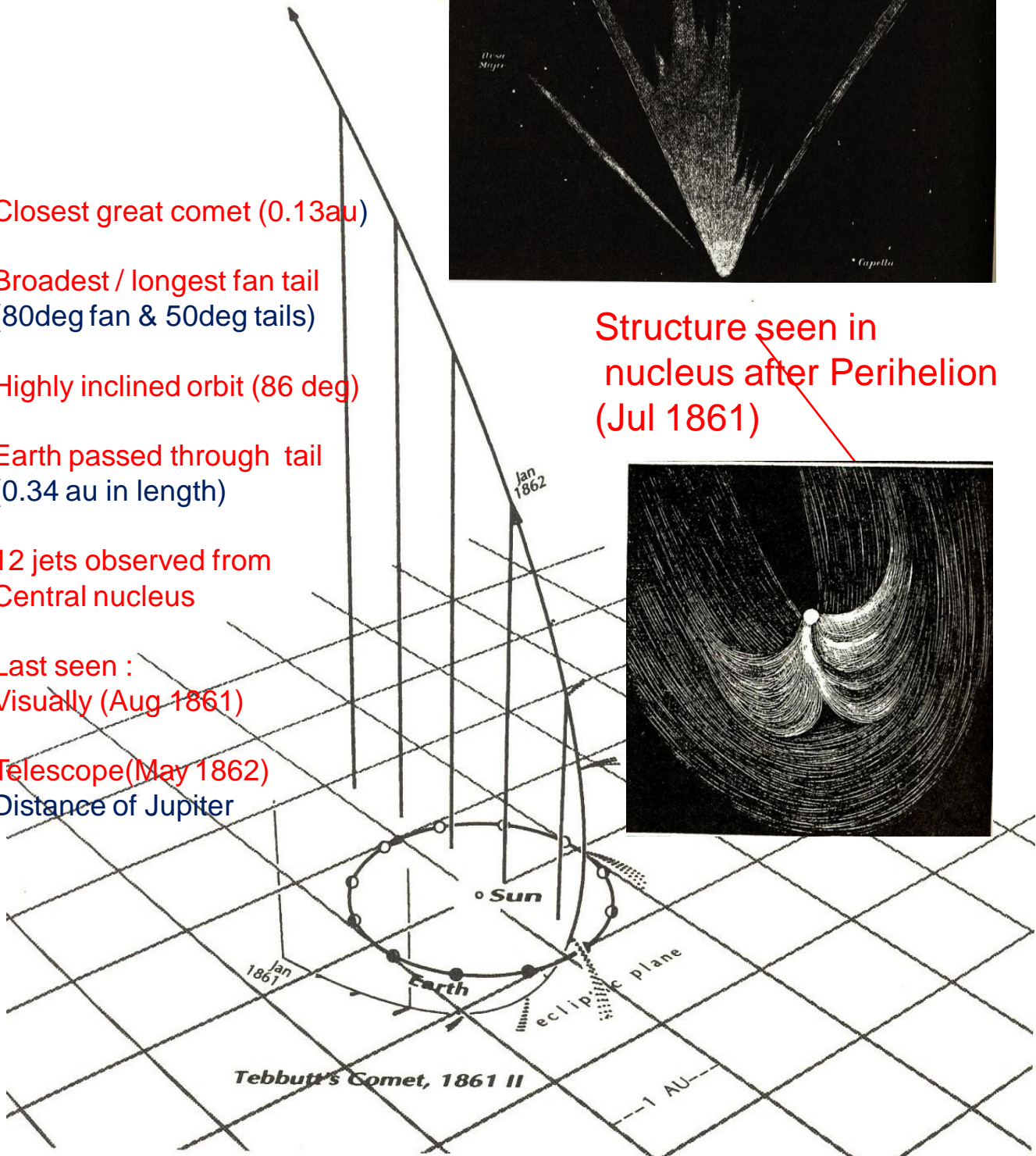
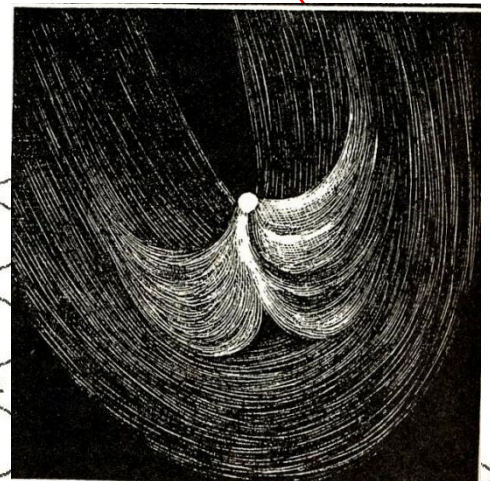
12 jets observed from
Central nucleus

Last seen :
Visually (Aug 1861)

Telescope (May 1862)
Distance of Jupiter



Structure seen in
nucleus after Perihelion
(Jul 1861)



Orbit of the great comet of 1861.

Great Comets (16th – 19th Century)

(d) 1882



The Great Comet of 1882, photographed by Sir David Gill.

Brightest comet in over 700 years
Near Sun

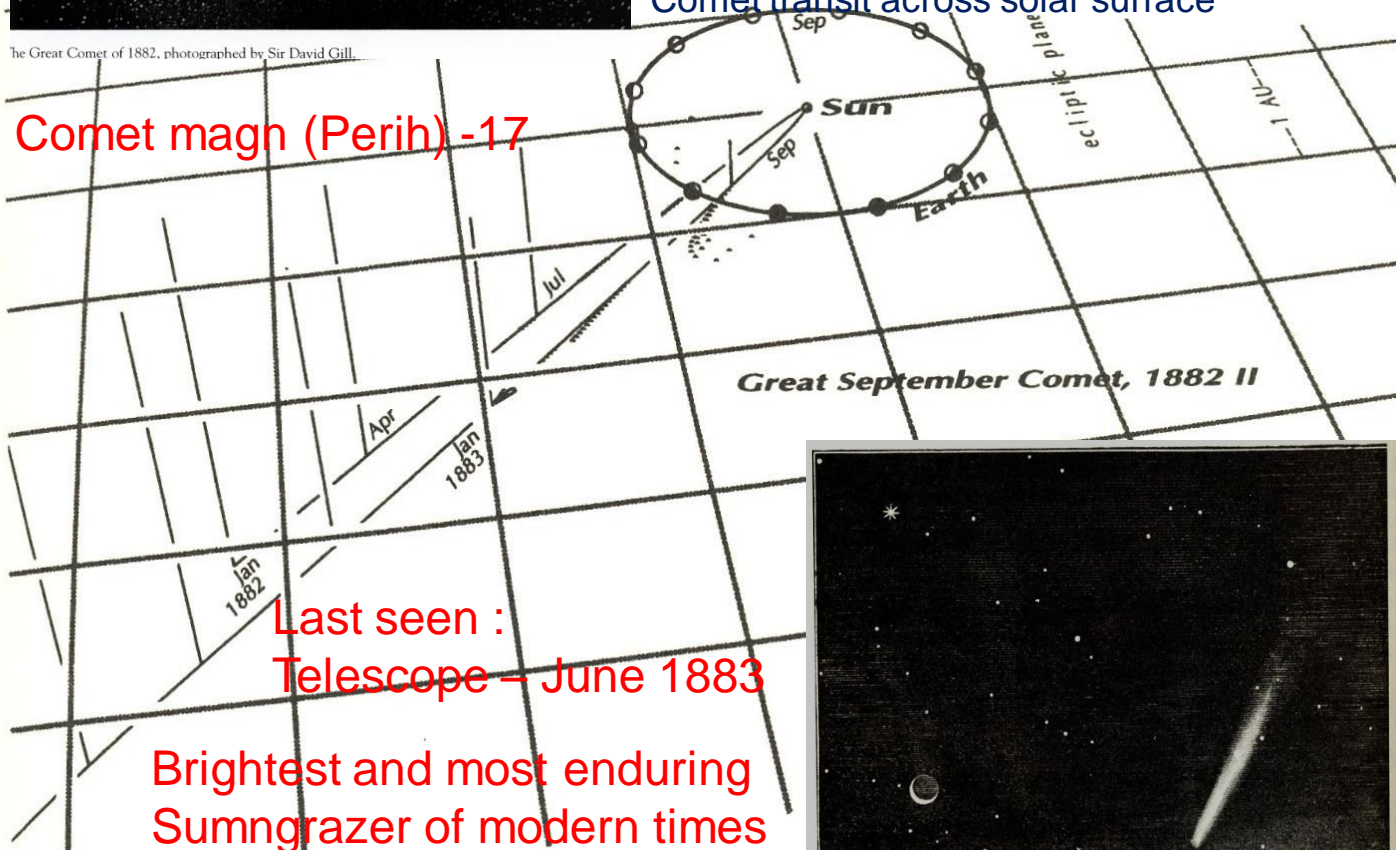
Bright as Venus (Sept 5 1882)

12deg from Sun (Sept 14)
seen in daylight

Perihelion – Sept 17 – only 464,000km
From solar surface – seen next to Sun's
Limb

Comet transit across solar surface

Comet magn (Perih) -17



Last seen :
Telescope – June 1883

Brightest and most enduring
Sungrazer of modern times



The Great September Comet of 1882 and the Moon.

Comet features :

Strange tail structure (above photo)

Split nuclei – day after perihelion

- Like a string of beads - 4 separate nuclei seen until Feb 1883- & tails

Great Comets (20th century– 2010)

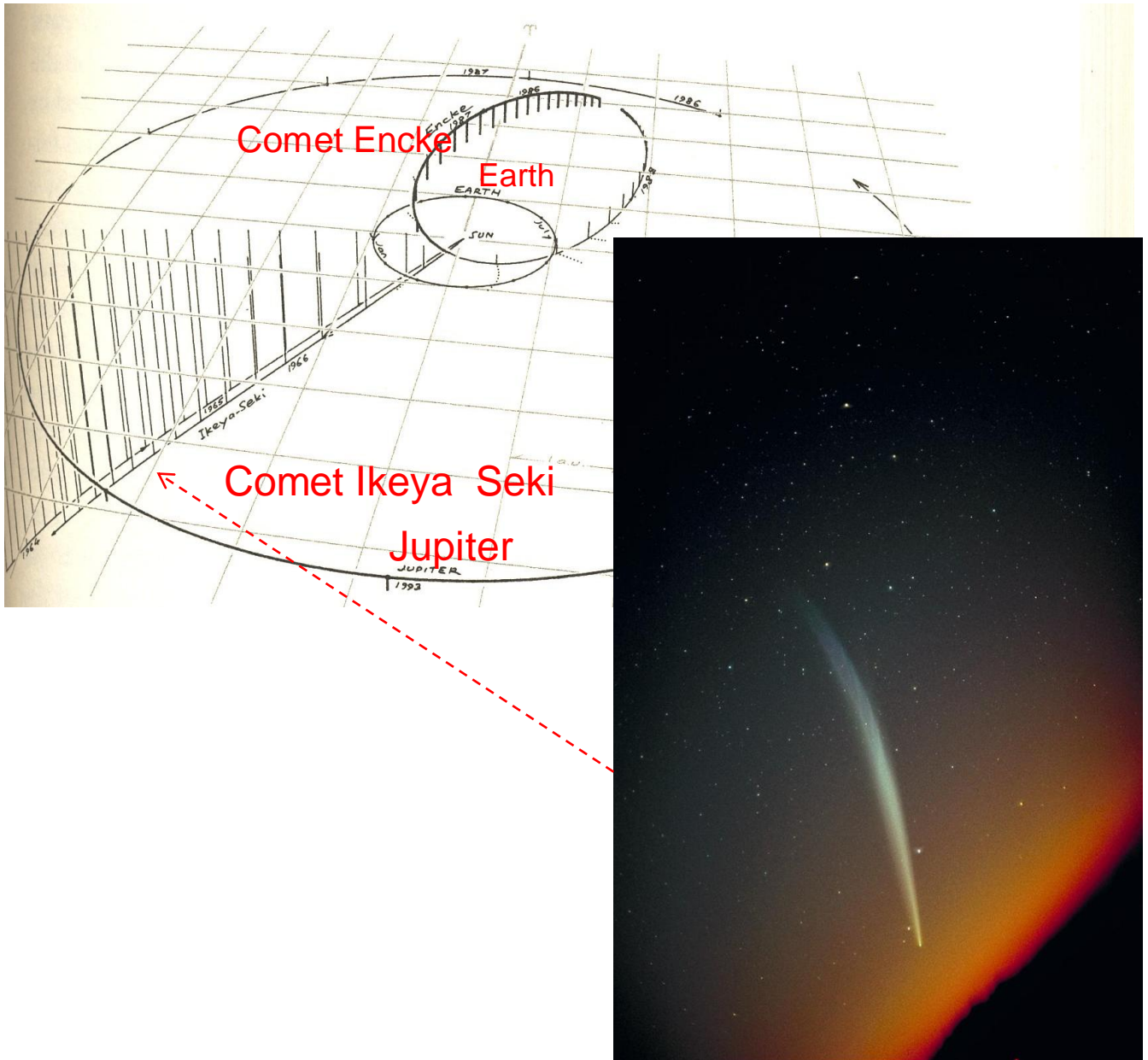
(c) Oct 1965 (Ikeya- Seki)

Greatest Kreuz sungrazer since 1882

(perihelion distance from Sun only 0.0008au ---Oct 21

Intense nucleus – magn -10 – split into 3 parts 30 min before Perihelion

Tail maximum length – over 60 deg (1.3 au)



Kitt Peak Observatory – Oct 29 1965

Great Comets (20th century– 2010) (d) 1976 (West)



Orbital period – 558,000 years

Perihelion on Feb 26th (0.2au)

Observed in daylight after perihelion

Best visibility – Mar 4th – closest to Earth

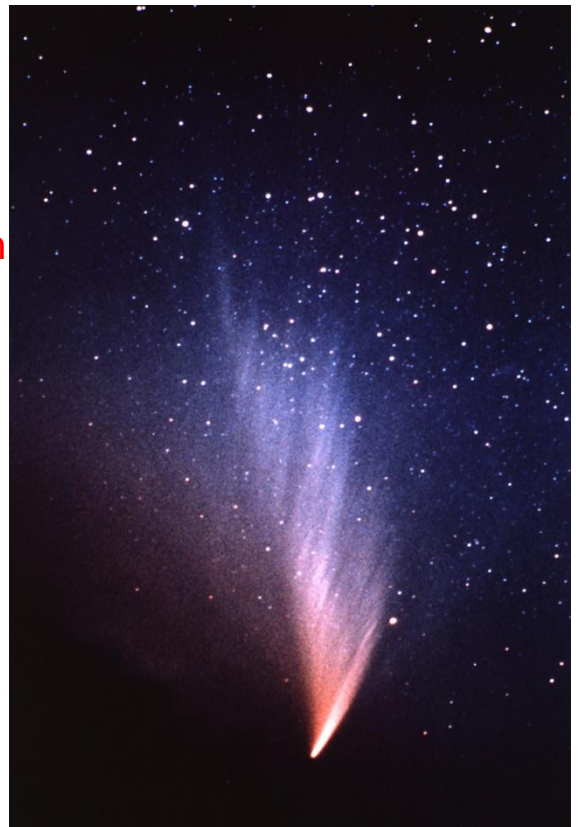
Gas / dust tails (30 deg) easily visible

Nucleus fragmentation(March) caused :

Dust tail structure

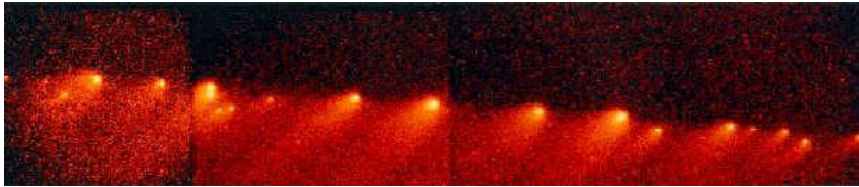
Unexpected brightness of comet

Last visual observation – Aug 25th
(magn 11)



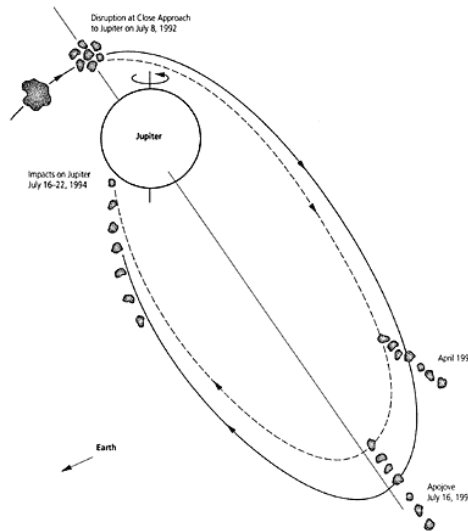
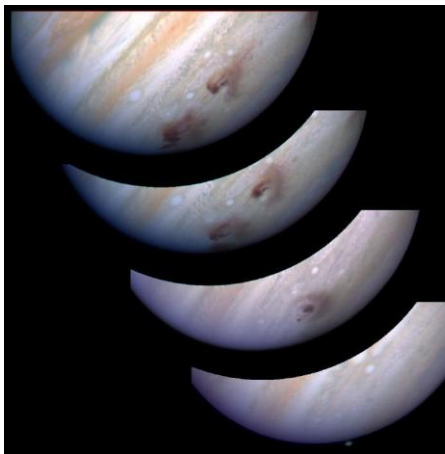
Great Comets (20th century– 2010)

(e) 1993 (Shoemaker-Levy 9 (SL9))

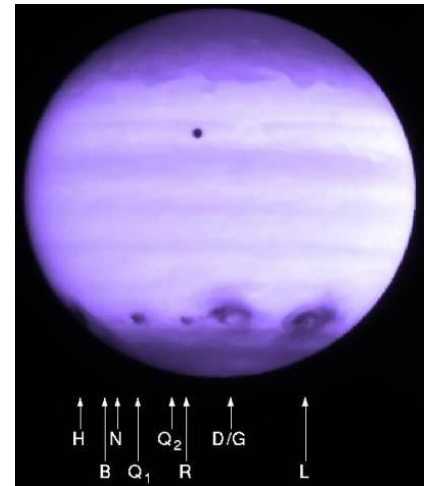


21 fragments from comet (HST image)

Impact sites (HST)



UV image of impact sites(HST)



Comet captured by Jupiter – fragmentation (Jul 1992) at perihelion

Discovered by Shoemaker/Levy – Mar 1993 - predicted collisions

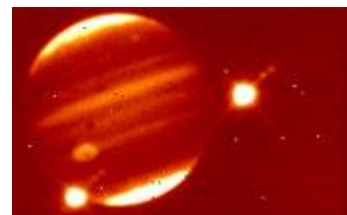
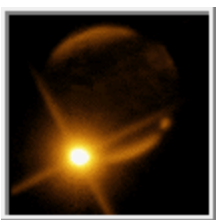
SL9 fragments – 100m to 1 km in size (original nucleus 5 km across)

Simultaneous spacecraft observations : Hubble / ROSAT/Galileo/Ulysses/Voyager2

21 impact sites observed – Jul 1994 – typically 12000km across (fragment G)

Energy released on collision – 6million tons of TNT(600x world`s nuclear arsenal)

Long term effects – visible scars for many months – global atmospheric changes



Great Comets (20th century– 2010) (e) 1997 (Hale Bopp)



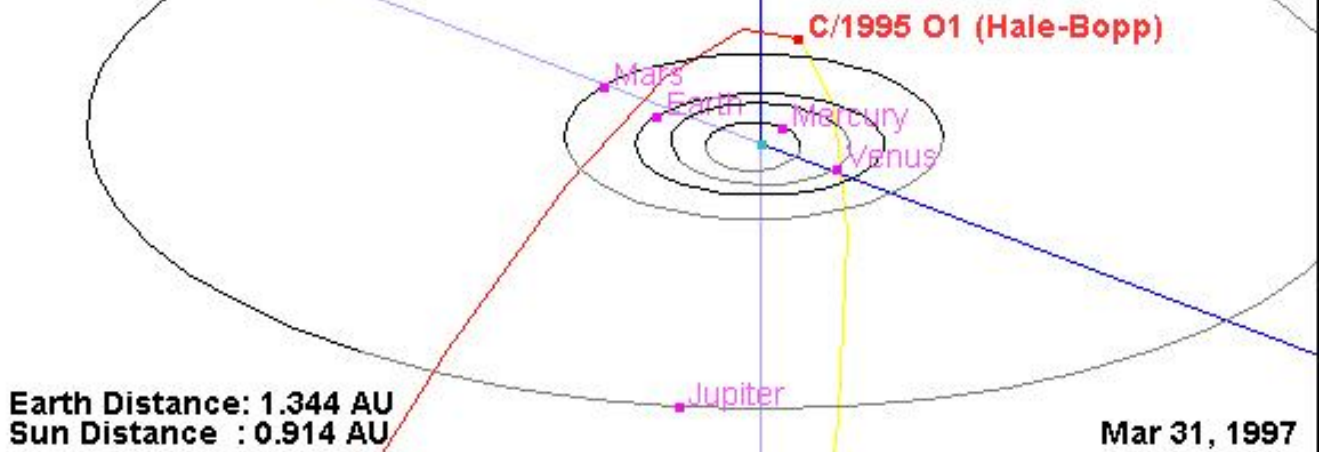
Feb 15th

C/1995 O1 (Hale-Bopp)



Mar 30th

Images taken by
S.Harding 10" Newtonian)



Comet of 20th Century :

Discovered Jul 23rd 1995 (7au)

Perihelion Apr 1st 1997 (0.914au)

- Almost as bright as Sirius

Dust tail of 40-45 deg

Still active in 2007 (25.7au)



Hale Bopp images (1)



Hale Bopp seen close to the North American Nebula (NGC7000)

March 4th 1997 (T & D Hallas)

165mm f/4 lens – 10min exposure

Hale Bopp images(2)



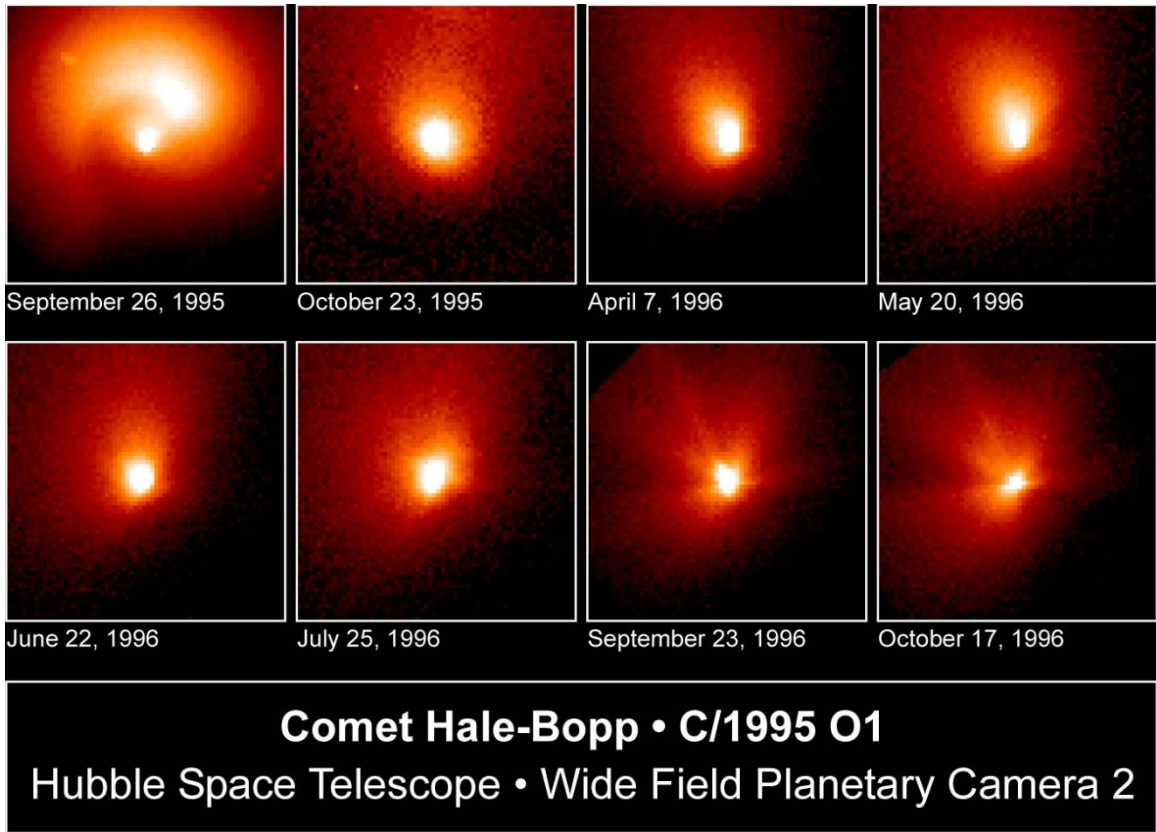
Hale Bopp seen close to Double Cluster
(open cluster M34 near comet`s head)

L.Tan 165mm lens 16min exp April 4th 1997

Great Comets (20th century– 2010)

(e) 1997 (Hale Bopp)

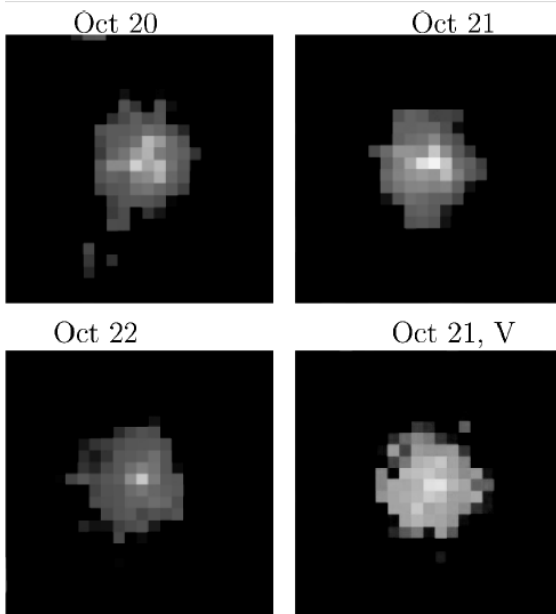
Beyond orbit of Jupiter



Images of the nucleus of Hale Bopp :

Outburst of dust beyond the orbit of Jupiter

1996 images show multiple jets connected with active nucleus vents



Images of Hale Bopp taken on Oct 20-21 2007 :

Diffuse coma of 180,000 km

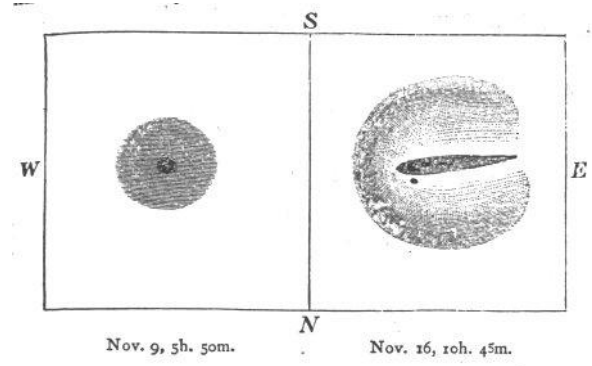
Magn +20 at distance of over 25AU

Comet activity detected (sublimation of CO)
(most distant activity detected in any comet)

(2.3m ANU telescope – Siding Spring)

Great Comets (20th century– 2010)

(g) 2007 (Comet Holmes)



Images of comet Holmes in 1892

17P/Holmes

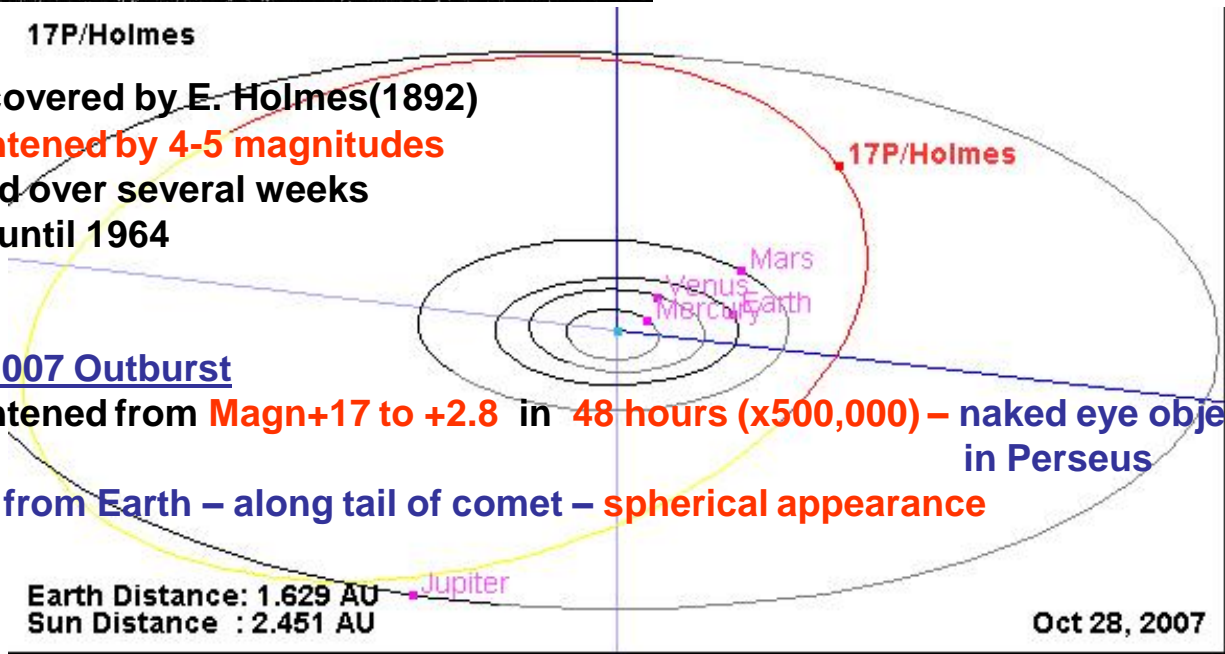
Discovered by E. Holmes(1892)

- Brightened by 4-5 magnitudes
- Faded over several weeks
- Lost until 1964

-Oct 2007 Outburst

-Brightened from **Magn+17 to +2.8** in **48 hours (x500,000)** – naked eye object in Perseus

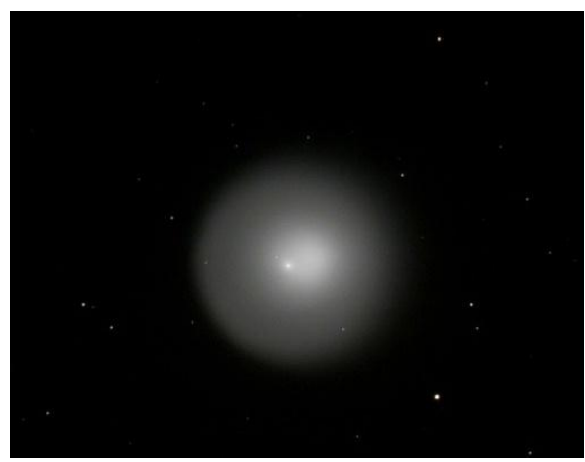
-View from Earth – along tail of comet – **spherical appearance**



Comet nucleus 3.4km



(28/10/07)



(30/10/07) S.Harding images(8" SCT)

Great Comets (20th century– 2010)

(f) 2007 (Comet Holmes)

Oct 2007- coma diameter of 13arc min (1million km at 2AU)



02/11/07 (8" SCT 60 sec)



06/11/07 (8" SCT 2 min)

Coma diameter –
x0.7 diameter of Sun

Outburst

– maybe caused by **sudden**
Venting of gas through
Nucleus

Comet fade to magn +5 /
Expanded to 2 deg (Feb2008)



wide field image(300mm lens) 12/11/07

(S.Harding images)



19 nov-07(10" SN)



28-nov-07(widefield 300mm lens)

(1)Spacecraft Missions to Comets

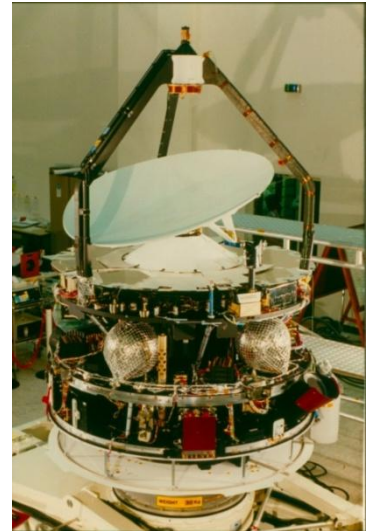
Giotto

First close-up images of a comet nucleus(March 1986)

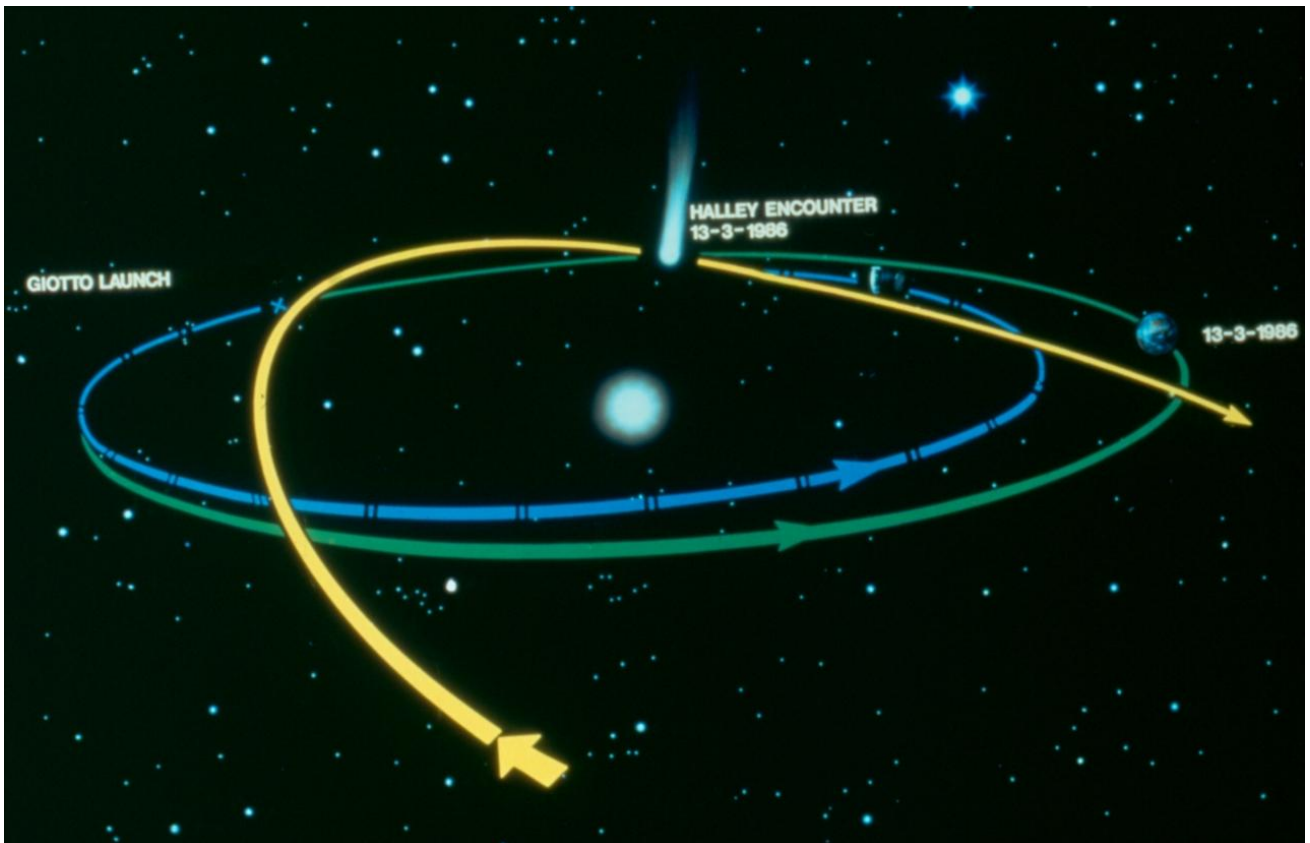


Key objectives:

- Determine composition of ices & Dust in Comet`s coma
- Measure comet`s gas production
- Rate
- Measure amount of dust around
- Comet & size/mass distribution
- investigate interaction of comet
- With solar wind particles

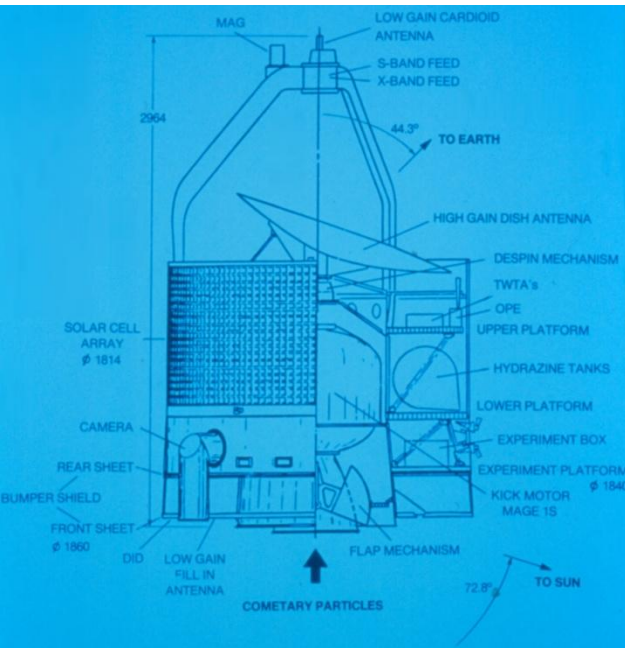


Launch 02/09/85 (Ariane 1)



First deep space mission using the Earth for gravitational assist

(1) Spacecraft Missions to Comets



Giotto Probe instruments:

Narrow angle multicolor camera – nucleus

3 mass spectrometers- gas/ dust composition

Plasma experiments – solar wind / charged particles

Energetic particle analyzer – electrons/protons/ alpha particles

Magnetometer – changes in magnetic field

Optical probe- brightness of nucleus

Radio science experiment solar wind environment

Giotto Events

March 13 1986 –

Halley crosses bow shock of solar wind
(solar particles slowed to subsonic speed in coma)

12,000 dust particle impacts occurred 2 hours before closest approach

Images transmitted up to 1372km from nucleus
(sharp increase in dust particles – Giotto passes through jet from nucleus)

7.6 sec before closest approach – comet sent spinning by impact with large (1 gm) dust particle

Giotto passed within 596km of nucleus



(1) Spacecraft Missions to Comets

Giotto Results

Results showed :

Nucleus – dark elongated body(15x10km) – **3 bright jets** – v dark

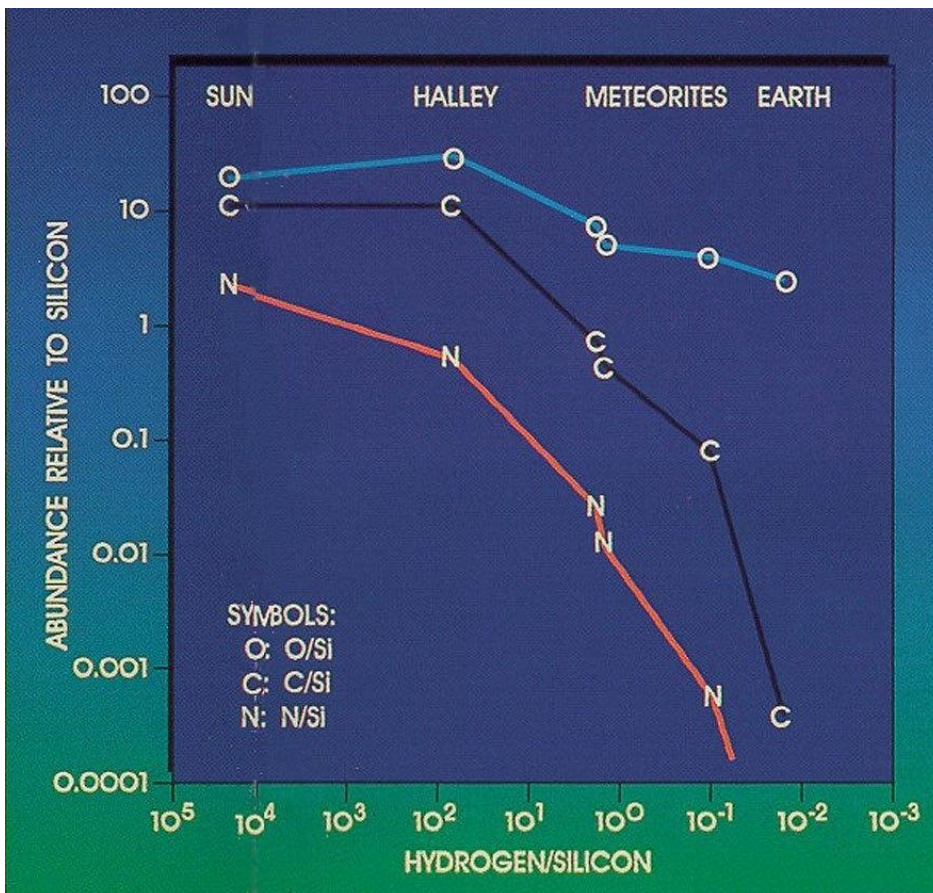
Jets – ejection rate of 3 tonnes/sec – **strange wobbling rotation of comet**

Dust grains – like smoke – particle mass about 40mg

Comet – formed 4.5 million yrs ago – **ices condensing onto interstellar dust grains**

Material ejected by comet (by vol) : 80% H₂O , 10% CO , 2.5% CO₂ + traces of NH₃ / CH₄

Abundances of elements relative to Silicon :



Dust particles :
contain CHON
Or
Na, Mg, Si ,Fe,Ca

All light elements
In comet (except N):

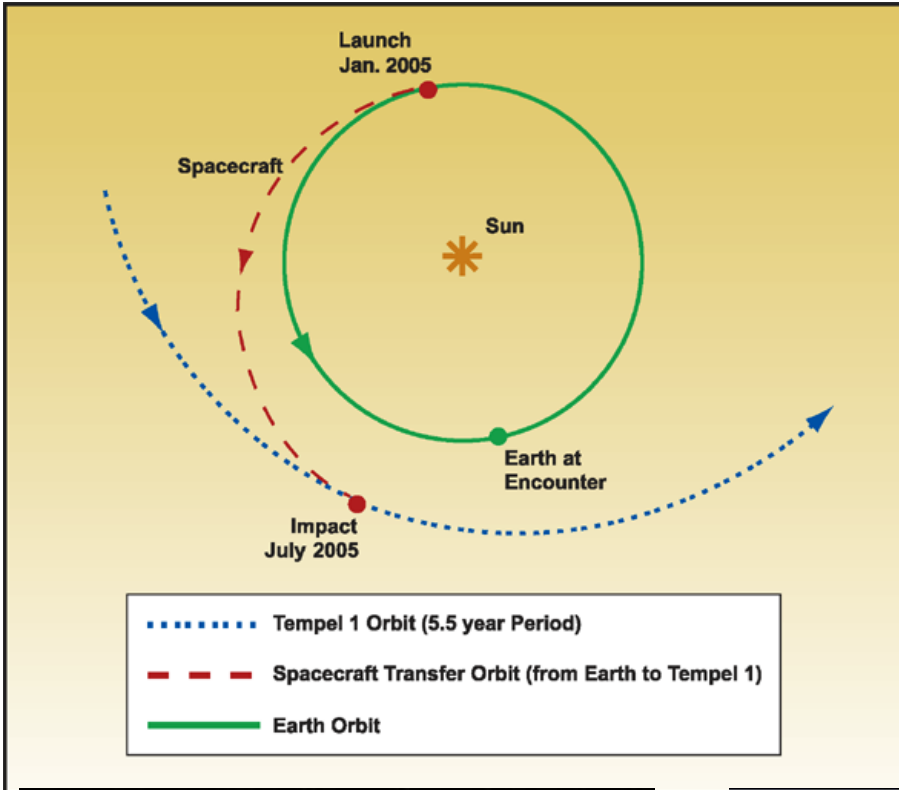
Same abundance
As in Sun –

Comet material
Is pristine and
Unprocessed

(2) Spacecraft Missions to Comets

Deep Impact Mission(2005)

Deep Impact Mission to Comet Tempel 1 – impactor sent to comet



Crater on comet :

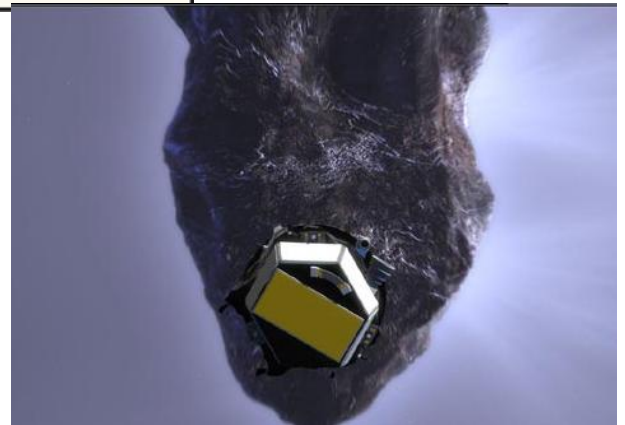
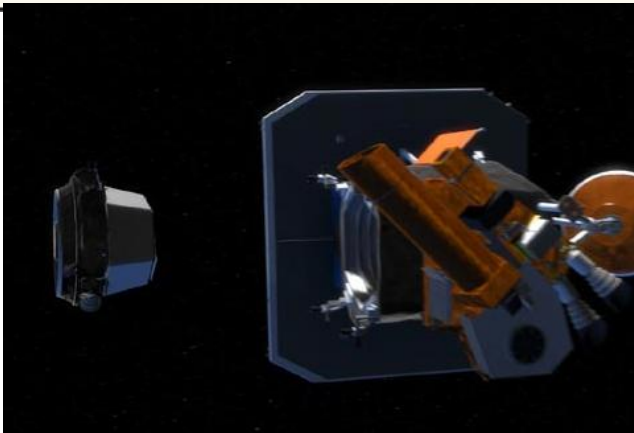
About size of stadium

Ice/ dust ejected & Observed from Spacecraft & Earth

Comet Tempel1:

Discovered in 1867

Orbital period 5.5 yrs



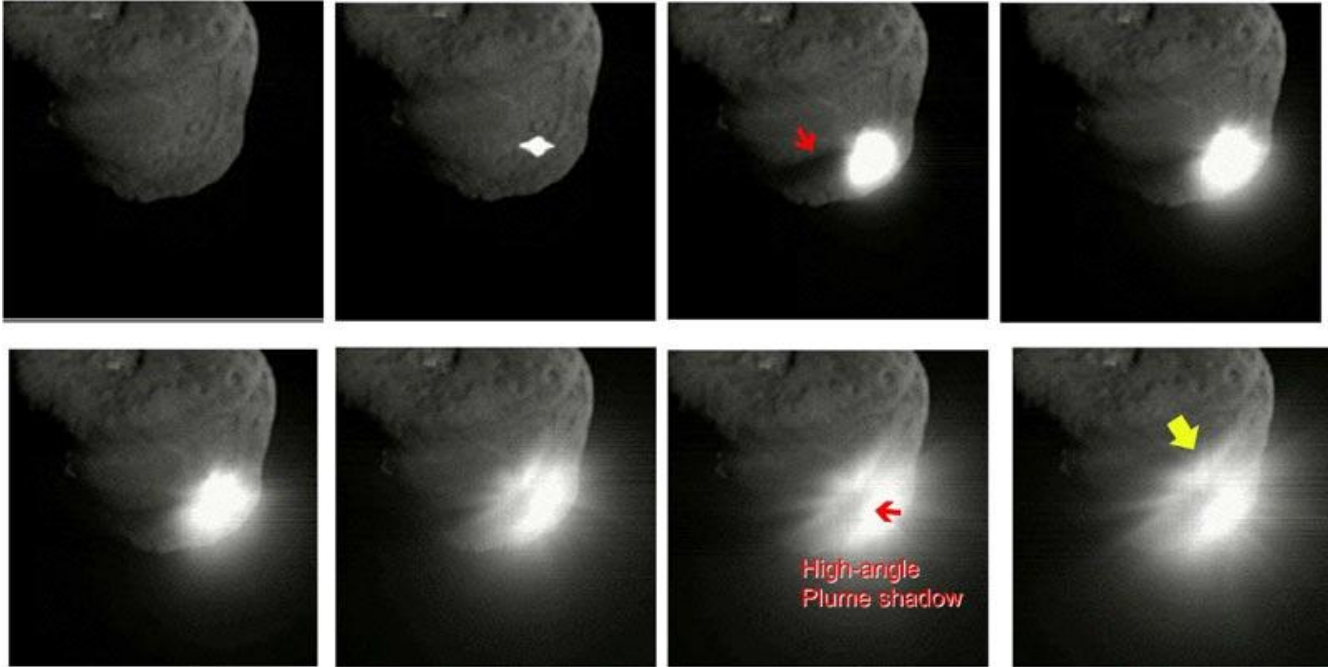
Release of impactor from Deep Impactor Spacecraft

Flyby spacecraft - 500km from comet :

observation of Impact / ejected material from crater

Structure / composition of crater interior

(2) Spacecraft Missions to Comets Deep Impact Mission (2005)



Deep Impact results showed for comet Tempel 1 showed :

a small amount of water ice on surface of comet

An abundance of organic material in comet's interior

The comet is a geologically active body – surface changing over time

(3) Spacecraft Mission to Comets

Stardust (2004) – Mission to Comet Wild2

Stardust mission Goal :

To return both particle samples from Comet Wild 2 and from interstellar dust.

Sample Return Mission :

Samples were returned to Earth for analysis:

Wild2 / Stardust

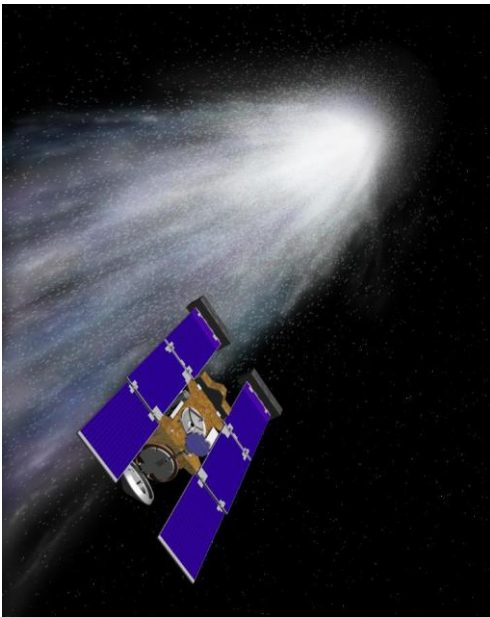


Image of Nucleus of Wild2



Interstellar dust / comet material captured in **aerogel**

Aerogel returned to Earth in **Sample Return Capsule**

Analysis of aerogel results :

Primitive organic matter found in aerogel – similar to **material in early solar system**

Polycyclic Aromatic Hydrocarbons (PAH) – found in sample – molecules common in interstellar space

Comet material may even contain **pre-solar system interstellar dust grains**

Visual / Photographic Observation of Comets

Steps to observe / image potentially bright comets :

(a) The orbit of the comet is usually known with great accuracy

- RA/Dec positions published on web sites
- Expected apparent magnitude

-(b) Wide field instrument can be used to scan region of Sky (eg binoculars)

- Bright comet will appear as diffuse object (few arc min in size)
- Beware of bright Deep sky object

-(c) Larger instrument will show structure of nucleus /coma

- Motion of comet apparent in field of view over short period
-

-(d) Astrophotography of comets

-**Large bright comets** – DSLR camera / 300mm lens

- (exposures up to 1-2min possibly to image faint tail structures)

-**Fainter comets** – prime focus imaging –DSLR/CCD camera

- shorter exposures (<1min) to avoid motion comet relative
- To background stars